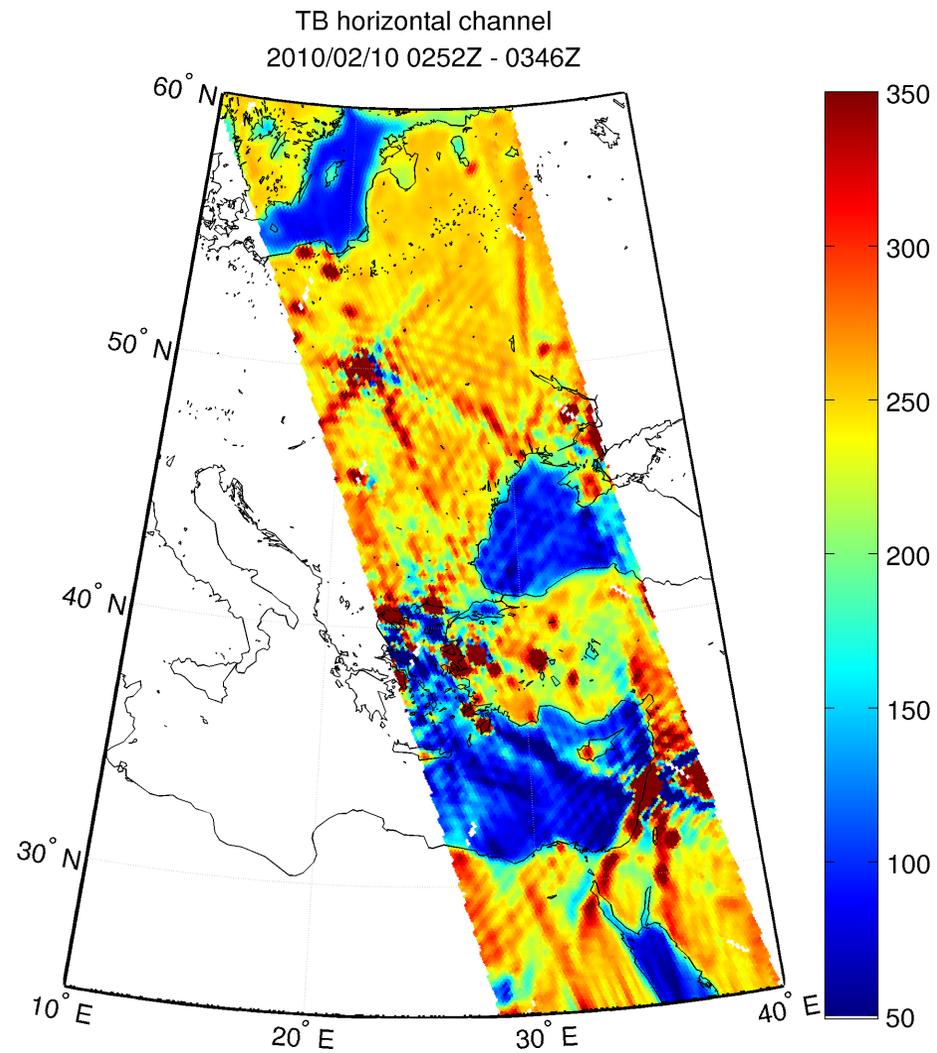


RFI in Airborne and SMOS Data. Aquarius/SAC-D Science Meeting, July 2010

**N. Skou, Jan Balling & Steen S. Kristensen
National Space Institute
Technical University of Denmark
ns@space.dtu.dk**

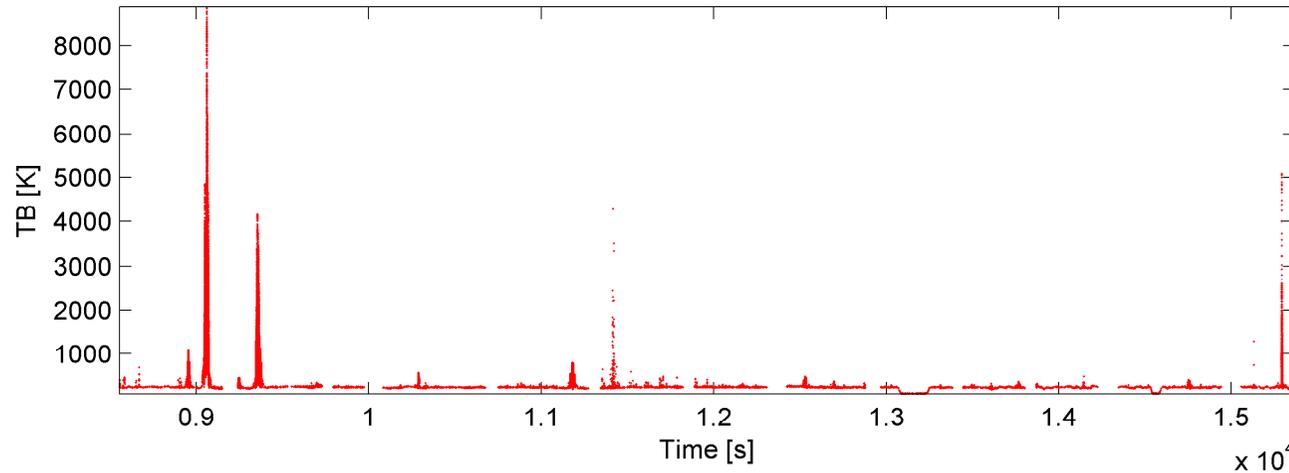


Airborne Data: Experience with Kurtosis

- In general only a few percent of data samples are flagged
- Less in Australia and North Europe
- More in South Europe
- Significant fraction close to airports and cities.
- **BUT: is kurtosis fully reliable?**
- We know from TB profiles and distribution plots that sometimes significant RFI goes undetected
- **This is truly annoying!**
- We know that kurtosis has a 50% duty cycle blind spot, and poor sensitivity in general for large duty cycles
- Can we improve the situation?

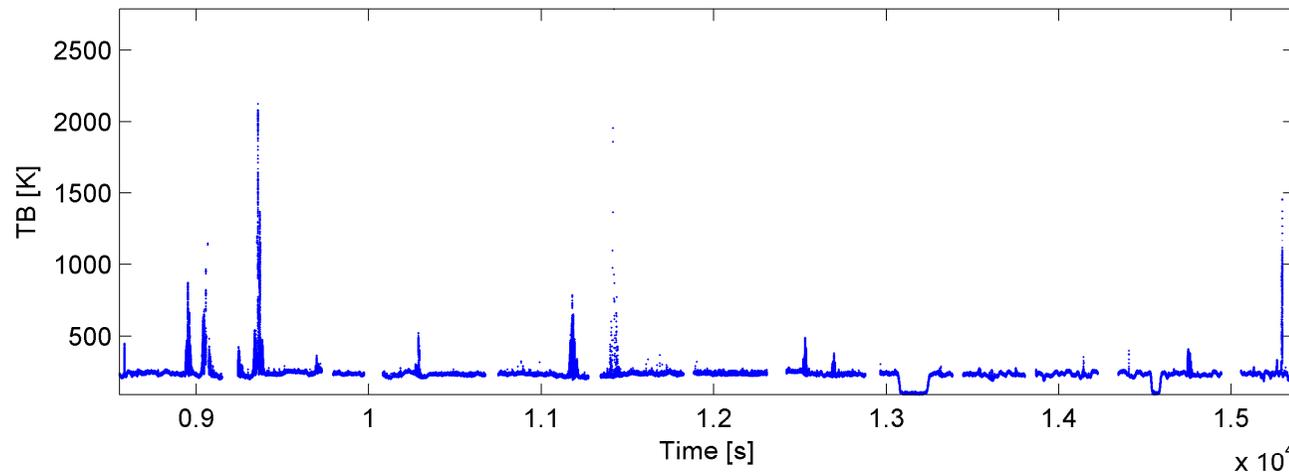
Standard method: 1 msec & +/- 4 std dev

V channel nadir, 20080419. kurt. limits +/- 4*stdDev(clean kurtosis)
Flagged samples



3.5%
flagged

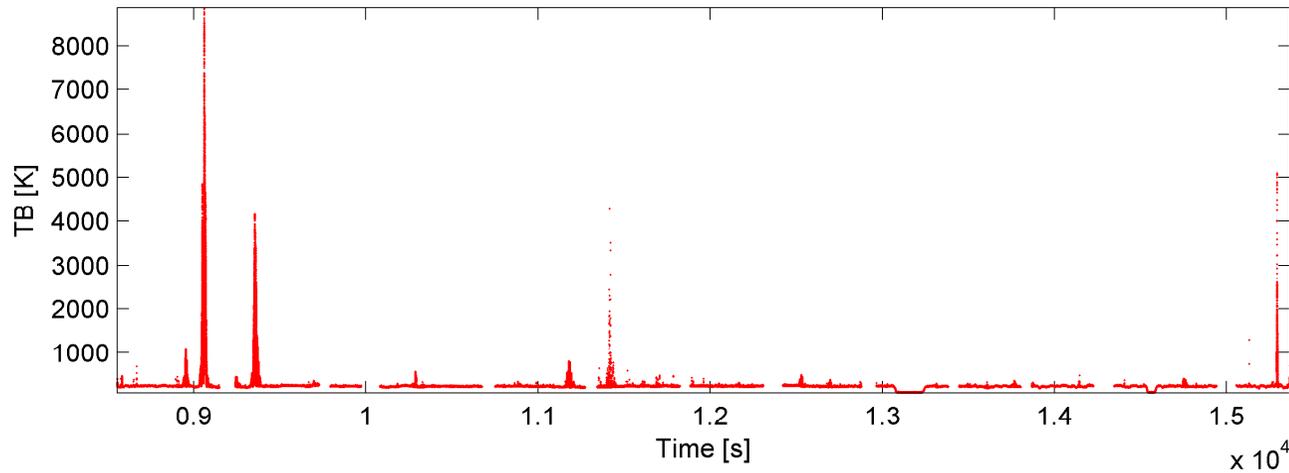
Clean samples



0.62%
> 350 K

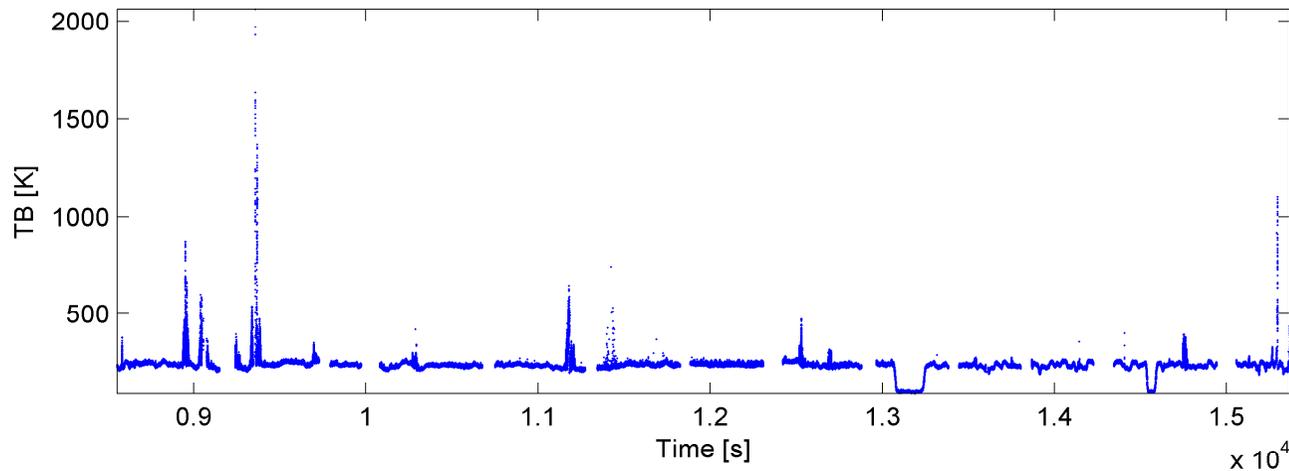
More sensitive method: 1 msec & +/- 0.5 std dev

V channel nadir, 20080419. kurt. limits +/- 0.5*stdDev(clean kurtosis)
Flagged samples



65%
flagged

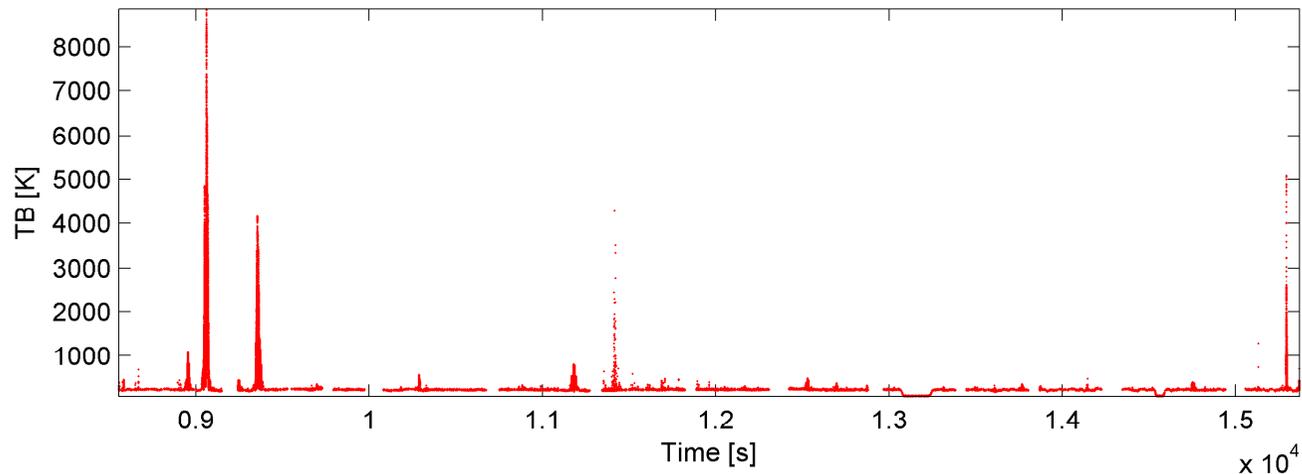
Clean samples



0.33%
> 350 K

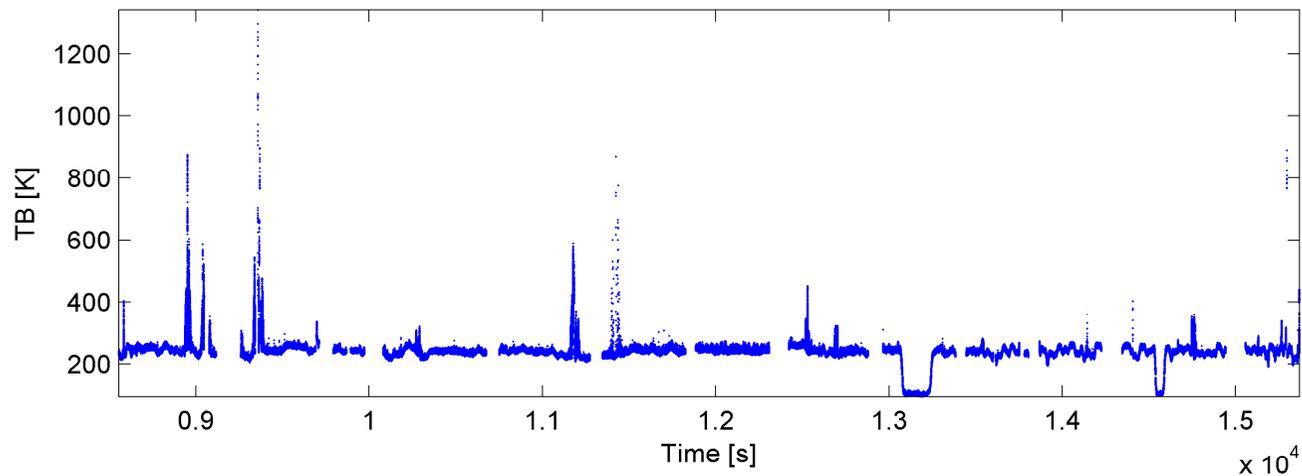
Discard flagged data + 30 neighbors

V channel nadir, 20080419. kurt. limits +/-4*stdDev,30 neighbouring samples blocked out
Flagged samples



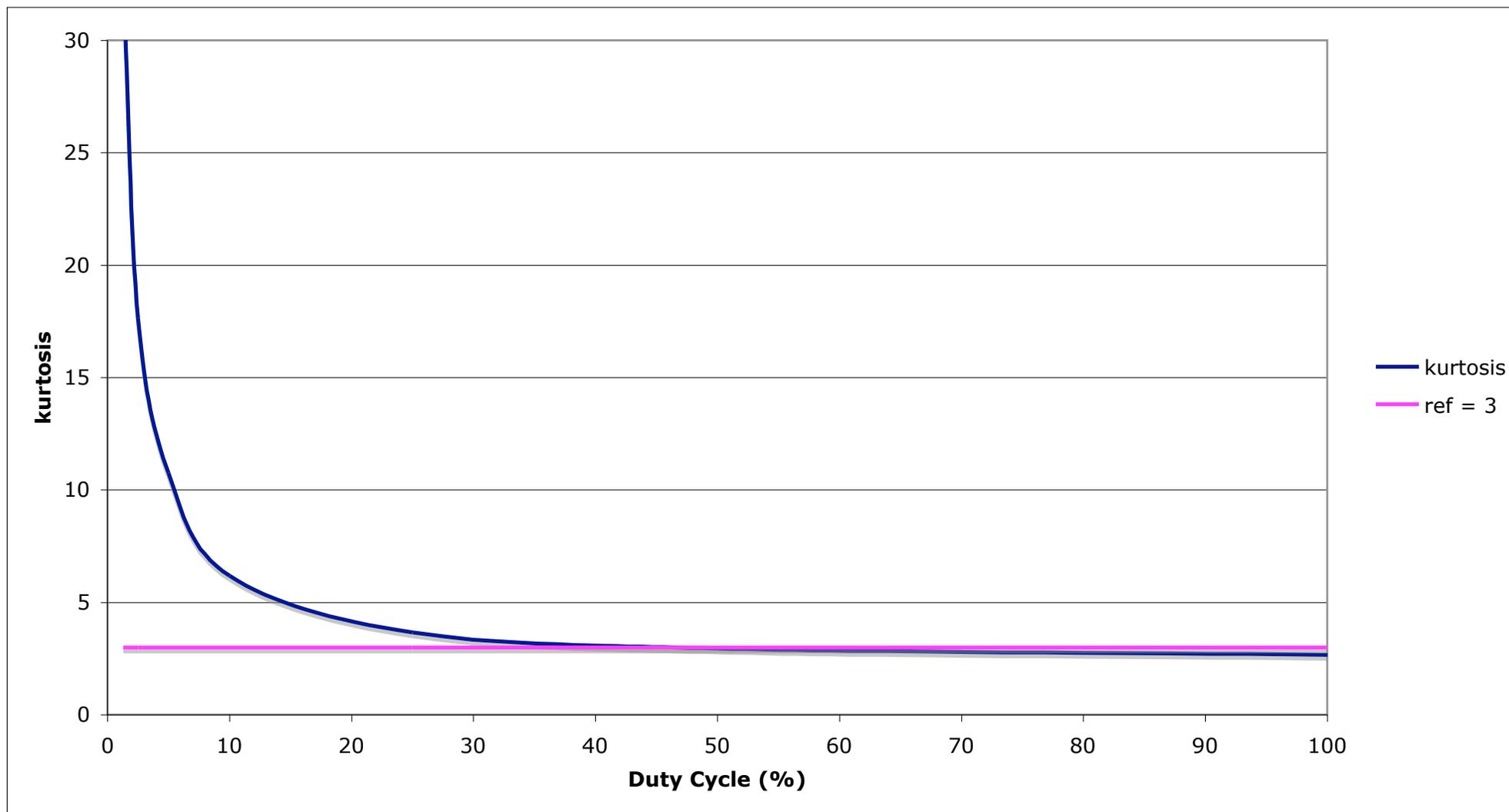
21%
flagged

Clean samples



0.3%
> 350 K

RFI contribution to TB constant = 100 K



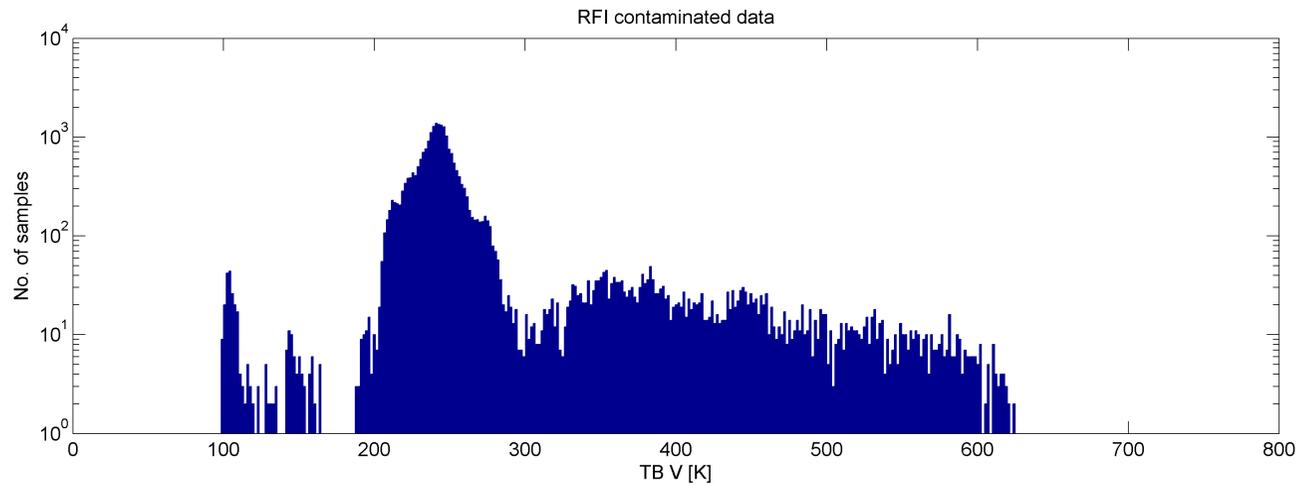
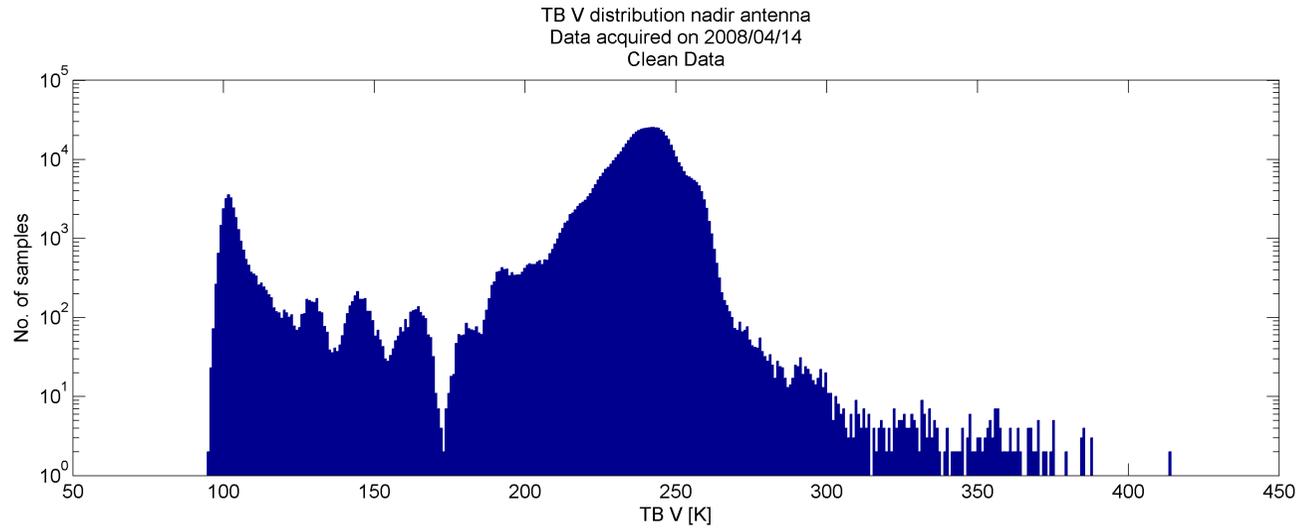
Kurtosis Method and “Blind Spot”

- Note that 50% duty cycle means 50% within radiometer calculation and integration window (1 msec in EMIRAD)
- Sensitivity is very dependent on duty cycle - and of RFI to TB ratio.
- In general: short, large, radar-like pulses very well detected by kurtosis.
- Large duty cycles much more difficult, i.e. missed detection of RFI!
- Possible remedy:
 - calculate kurtosis for different integration periods
 - evaluate together
 - several examples show no dramatic improvement!
- The problem is not the blind spot - but the poor detection of large duty cycles!

Polarimetric Data

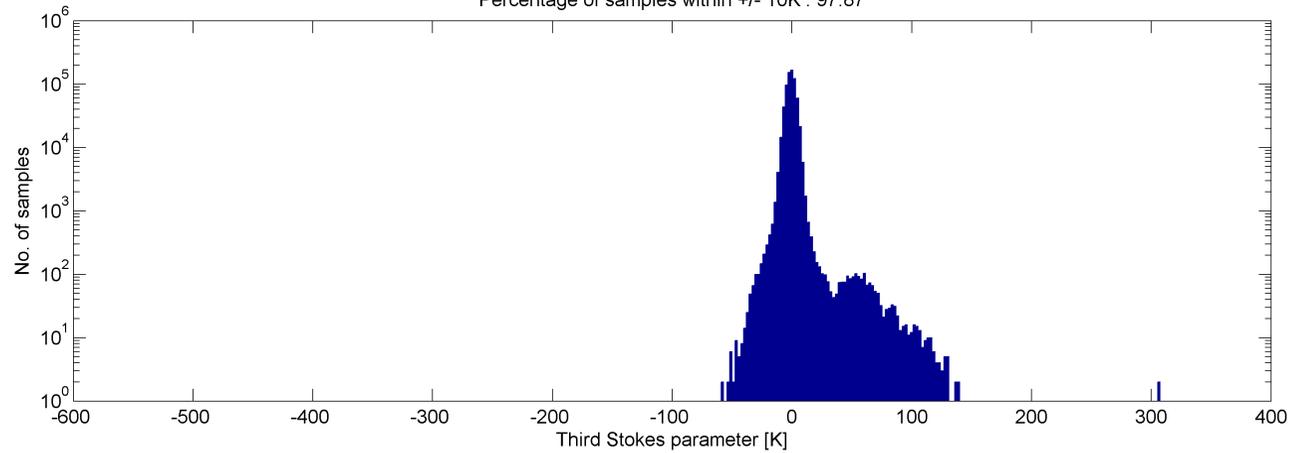
- **Natural targets have very small 3rd and 4th Stokes**
- **Linearly polarized RFI normally not aligned with H and V of our instrument, hence we get 3rd Stokes**
- **Many surveillance radars use circular polarization, hence we get 4th Stokes**
- **Experience with EMIRAD shows that often Kurtosis flagged data has significant signals in 3rd and 4th Stokes, but not always. The opposite can also be the case.**
- **Subject for further investigations**
- **Anyway, looking for signals in the 3rd and 4th Stokes channels of SMOS can be an important method for RFI detection**

TBV Distribution

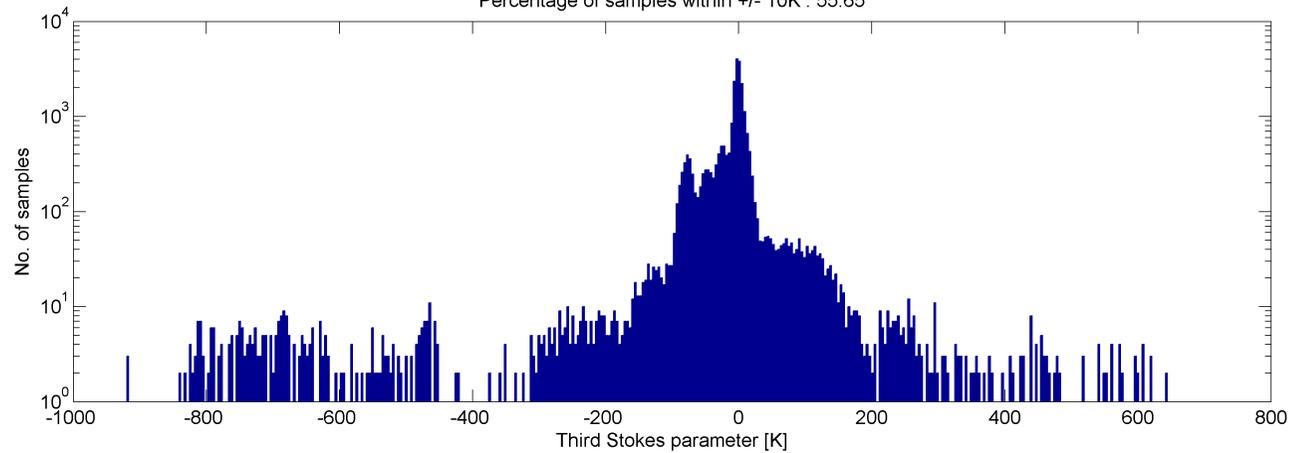


3rd Stokes Distribution

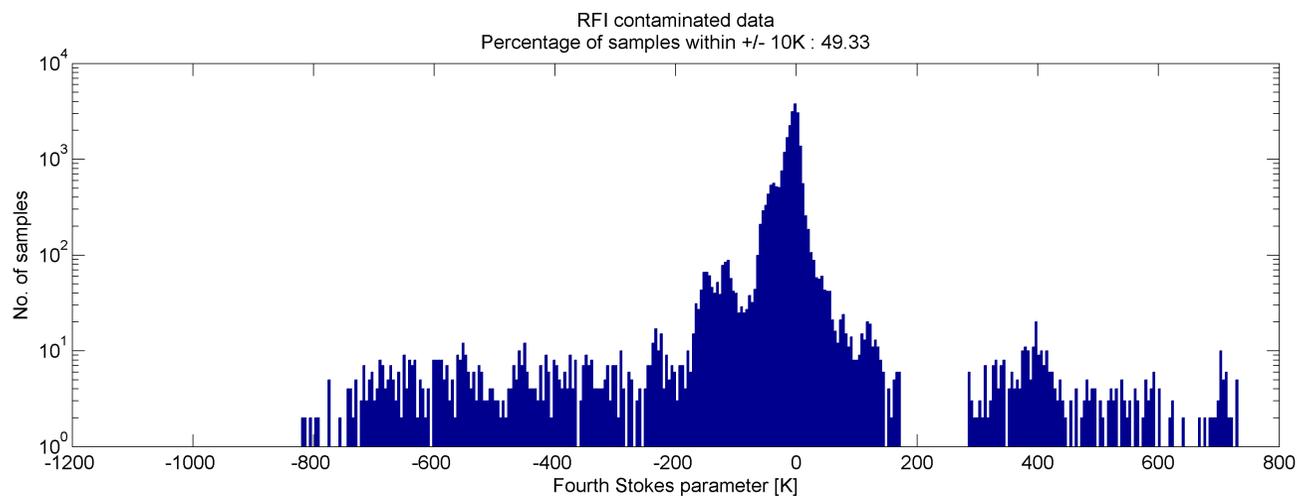
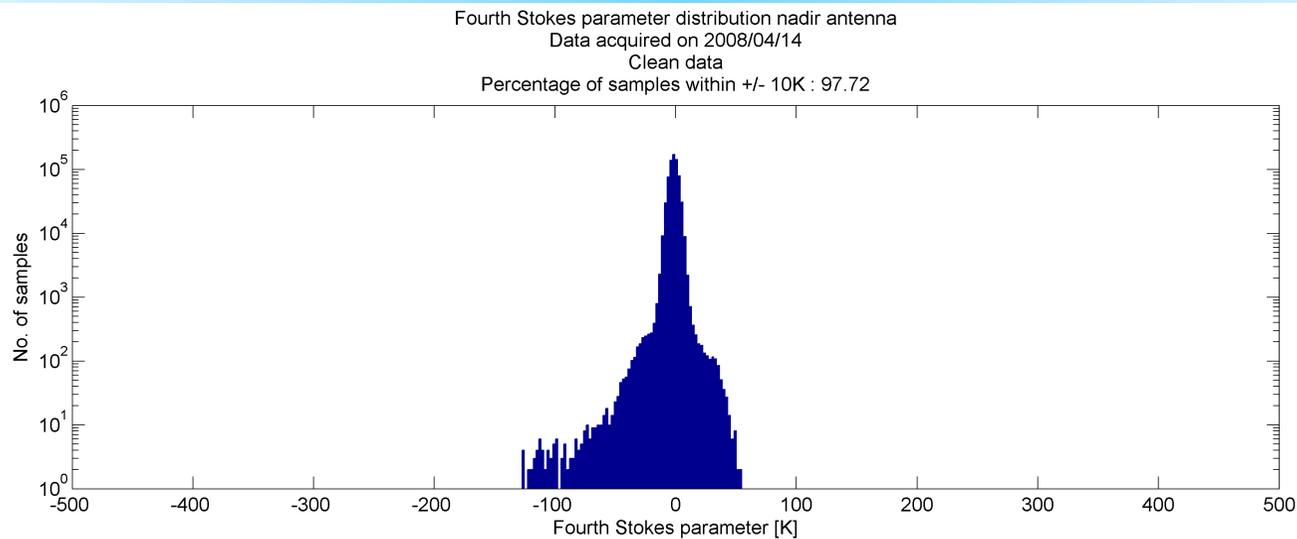
Third Stokes parameter distribution nadir antenna
Data acquired on 2008/04/14
Clean data
Percentage of samples within +/- 10K : 97.87



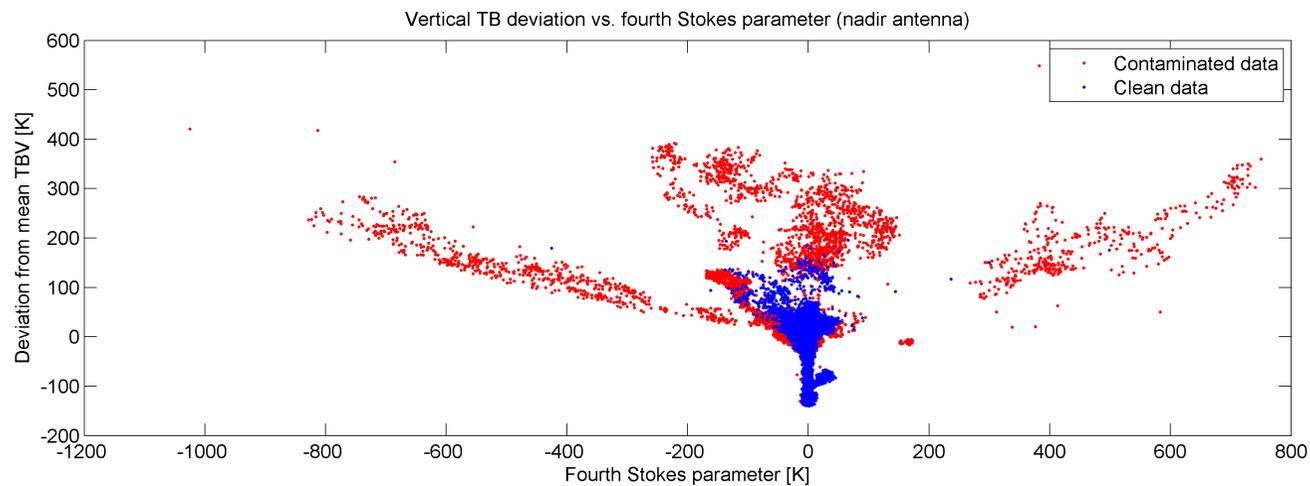
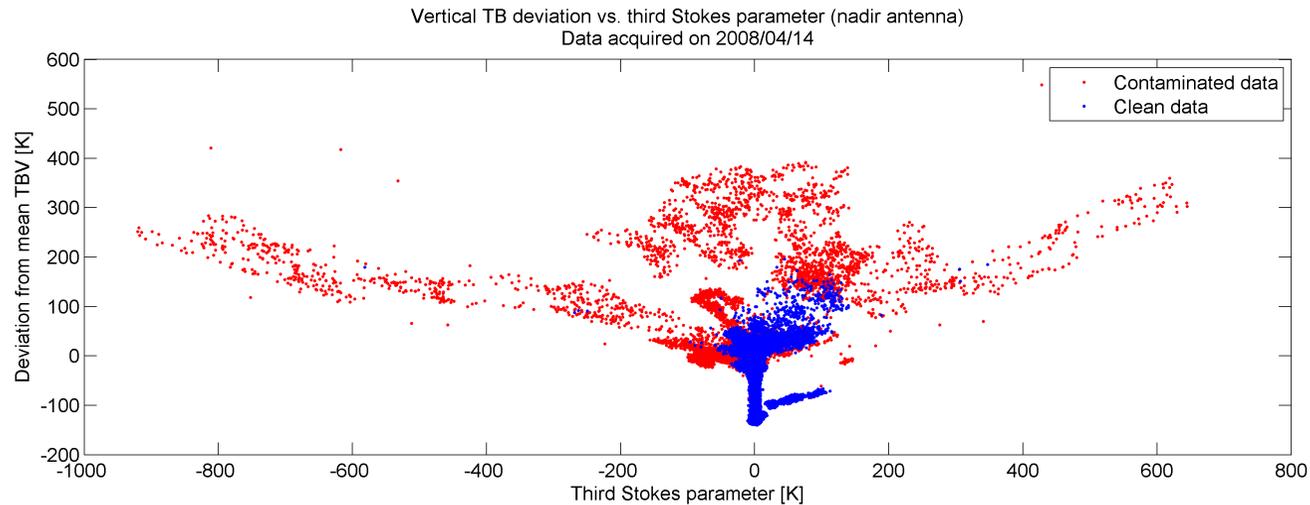
RFI contaminated data
Percentage of samples within +/- 10K : 55.65



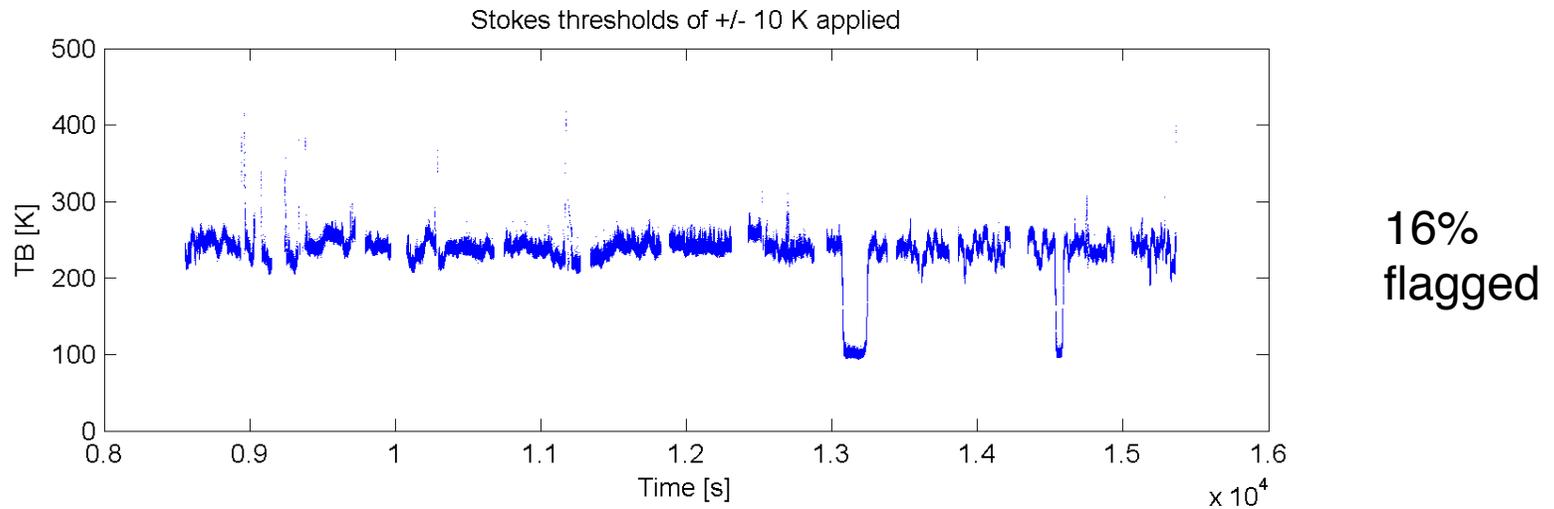
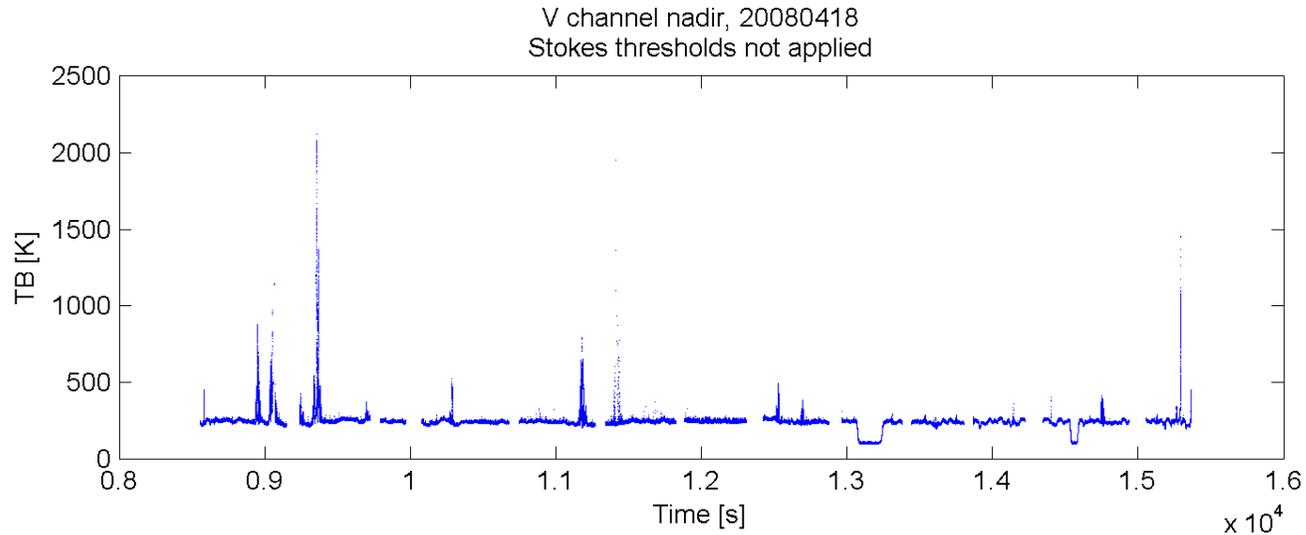
4th Stokes Distribution



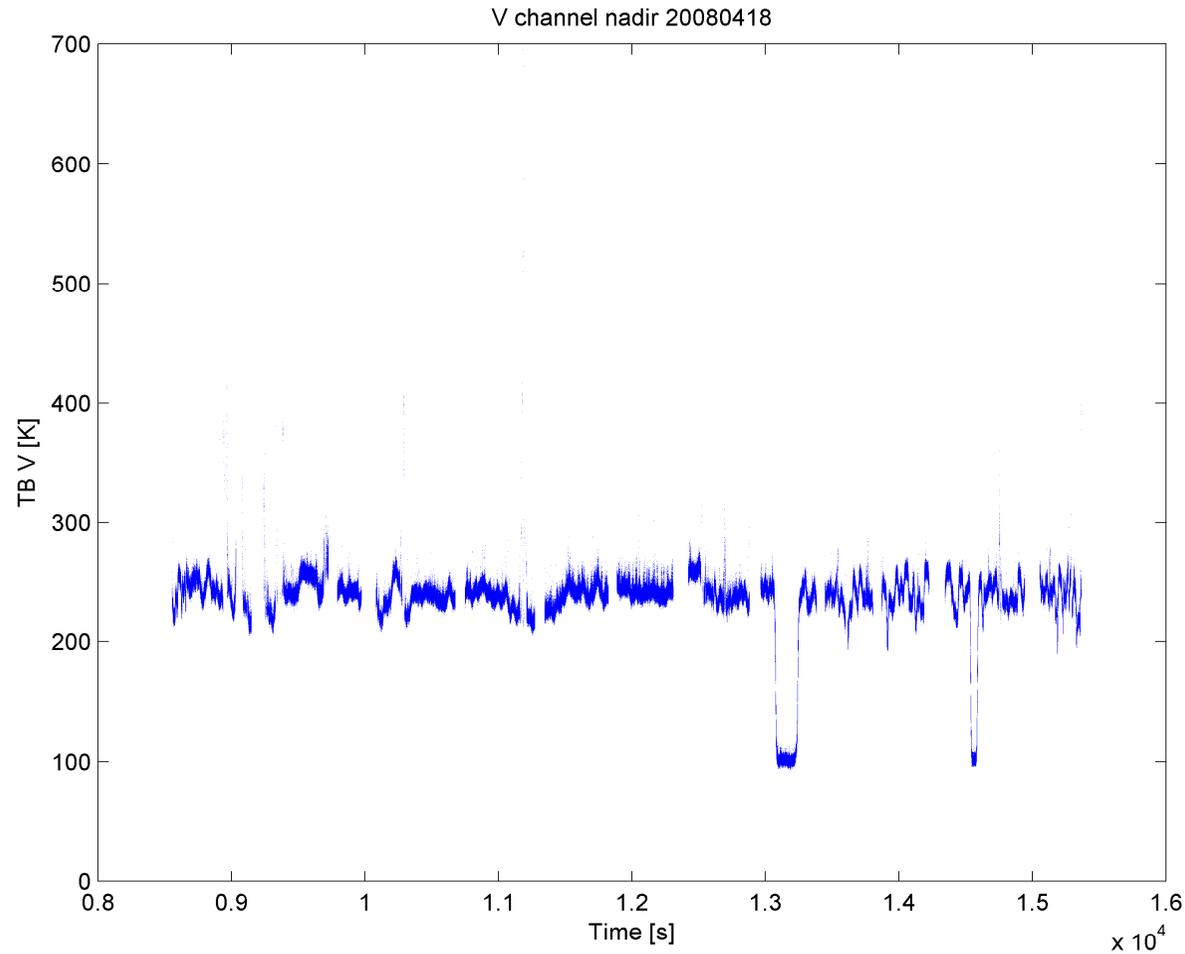
TBV vs. 3rd & 4th Stokes (EMIRAD, München)



Kurtosis AND ± 10 K threshold for 3rd and 4th Stokes



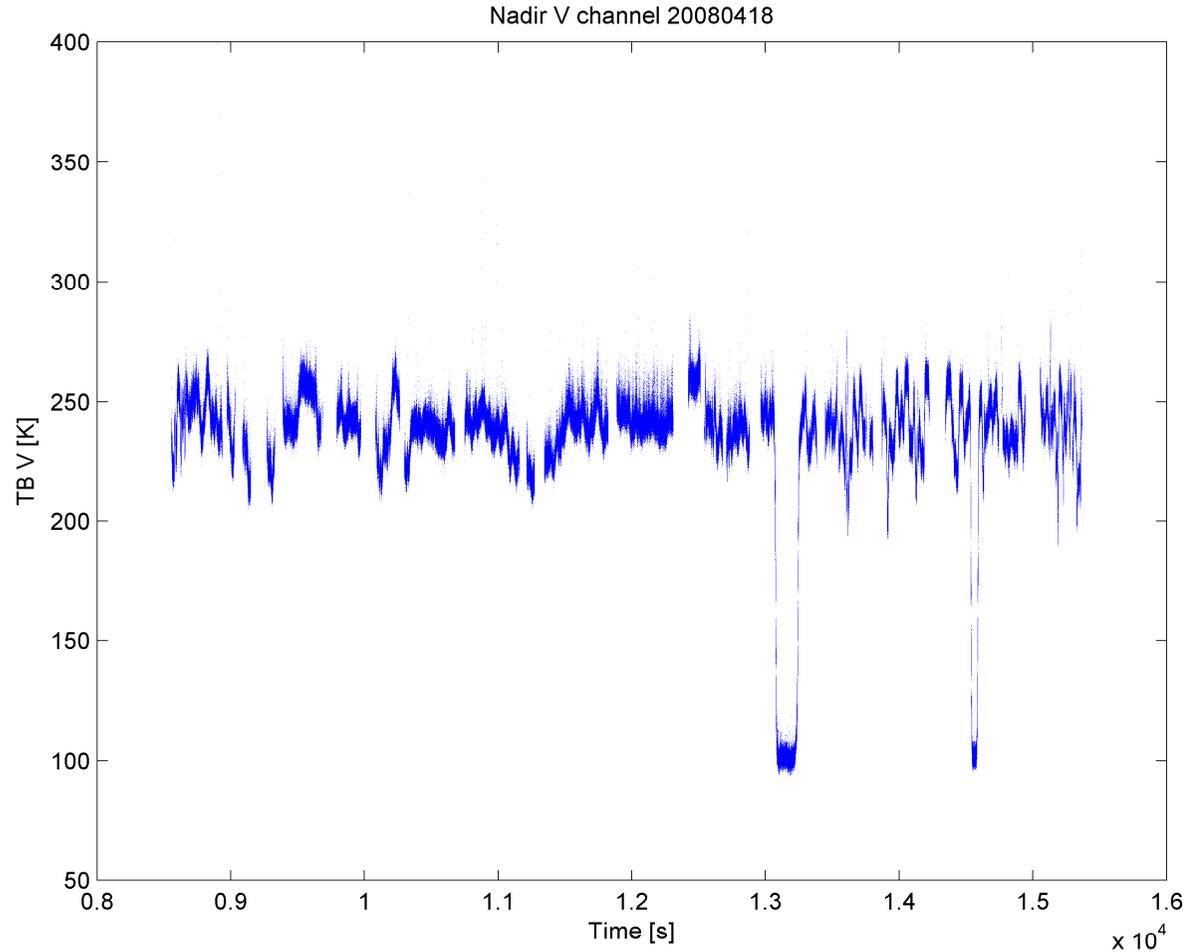
± 10 K Threshold for 3rd and 4th Stokes Applied



15% flagged

0.01% > 300 K

Thresholds for Std. Dev. of 3rd and 4th Stokes



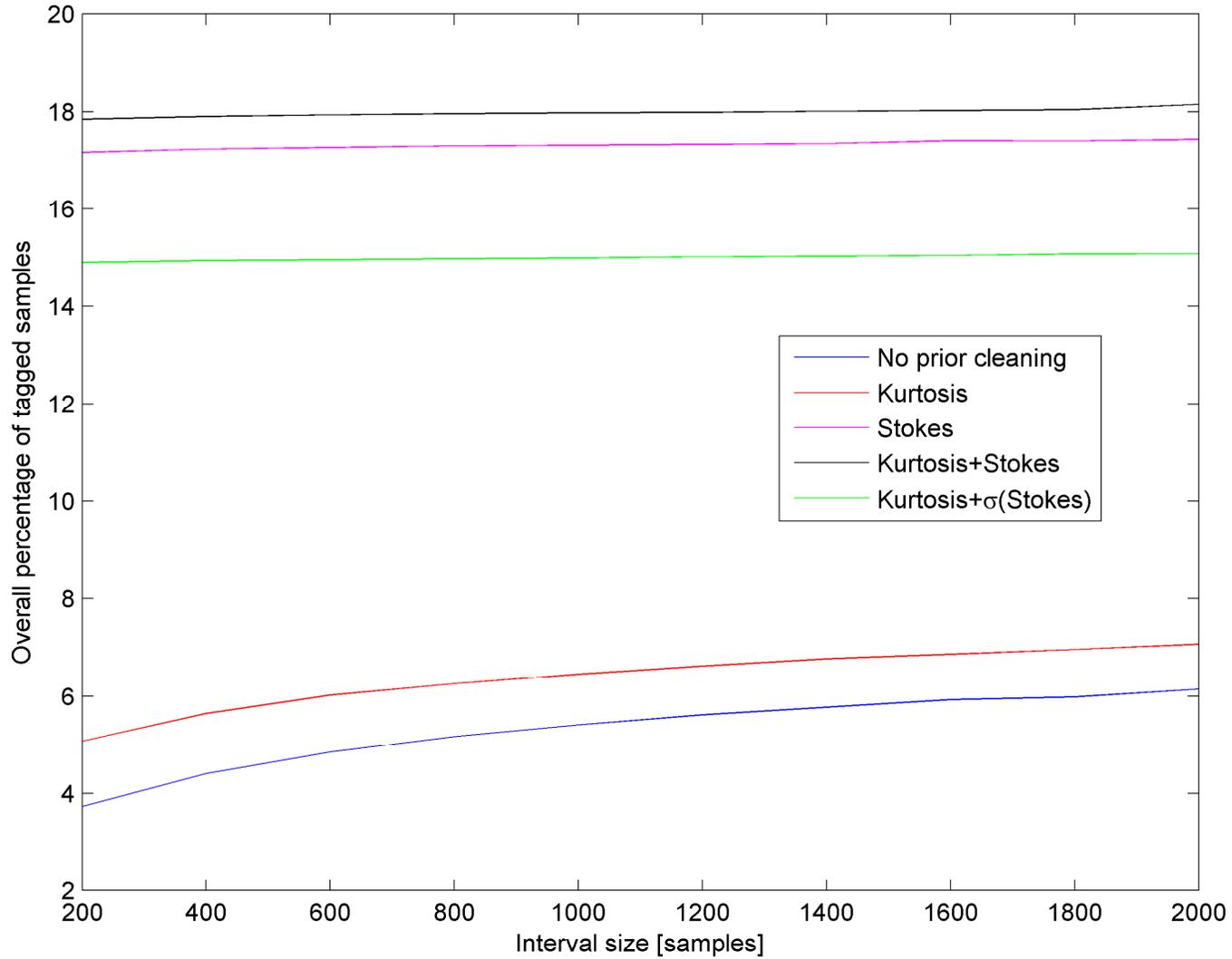
14% flagged

0.001% > 300 K

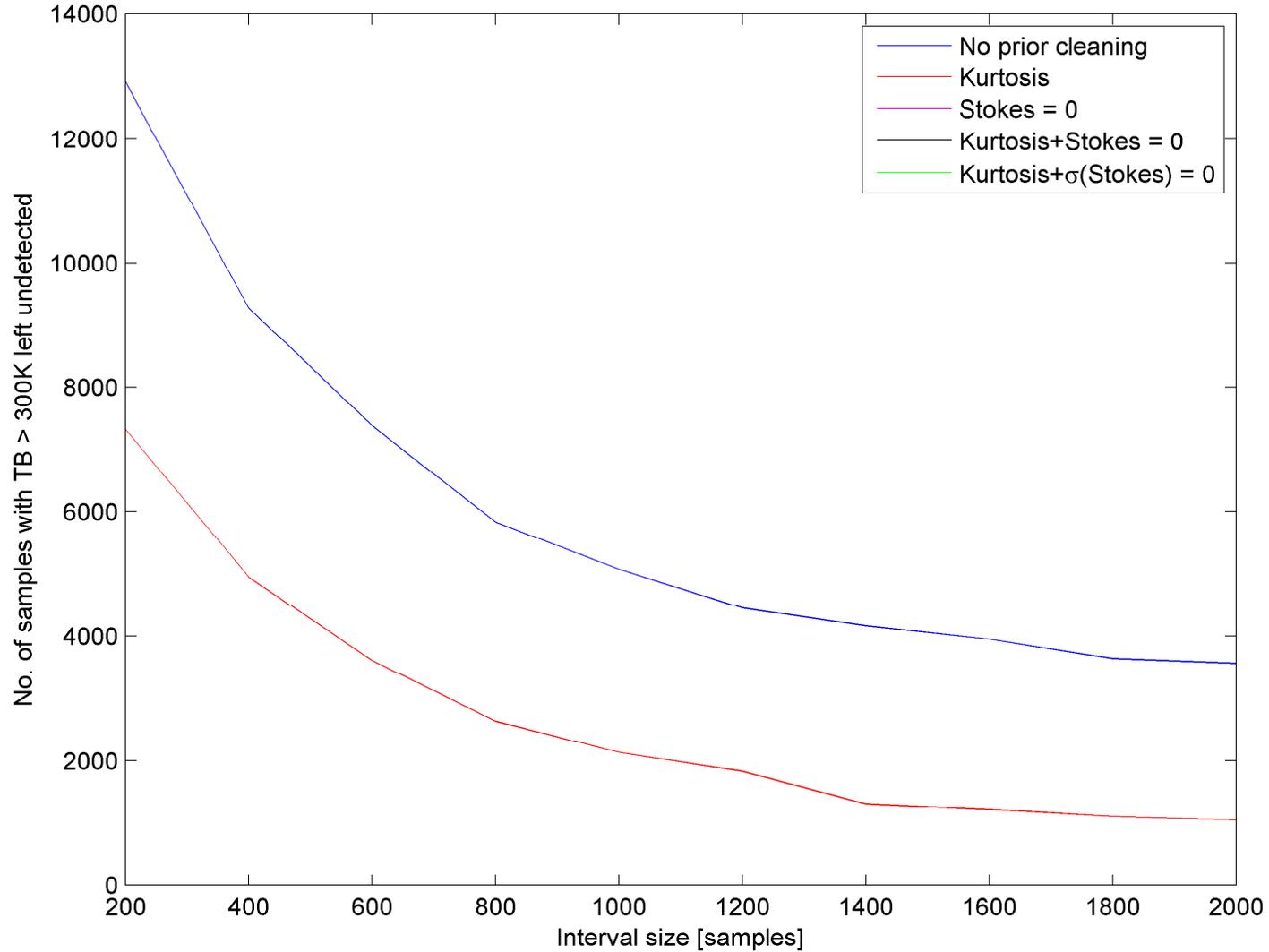
Glitch Detector

- Developed by Ruf et al.
 - Each sample compared with mean of neighboring samples
 - Discarded if it exceeds by a threshold
 - No. of neighboring samples is parameter
-
- **Glitch detector cleans up very well after polarimetry has been used!**

V channel of nadir antenna, 20080418



V channel of nadir antenna, 20080418

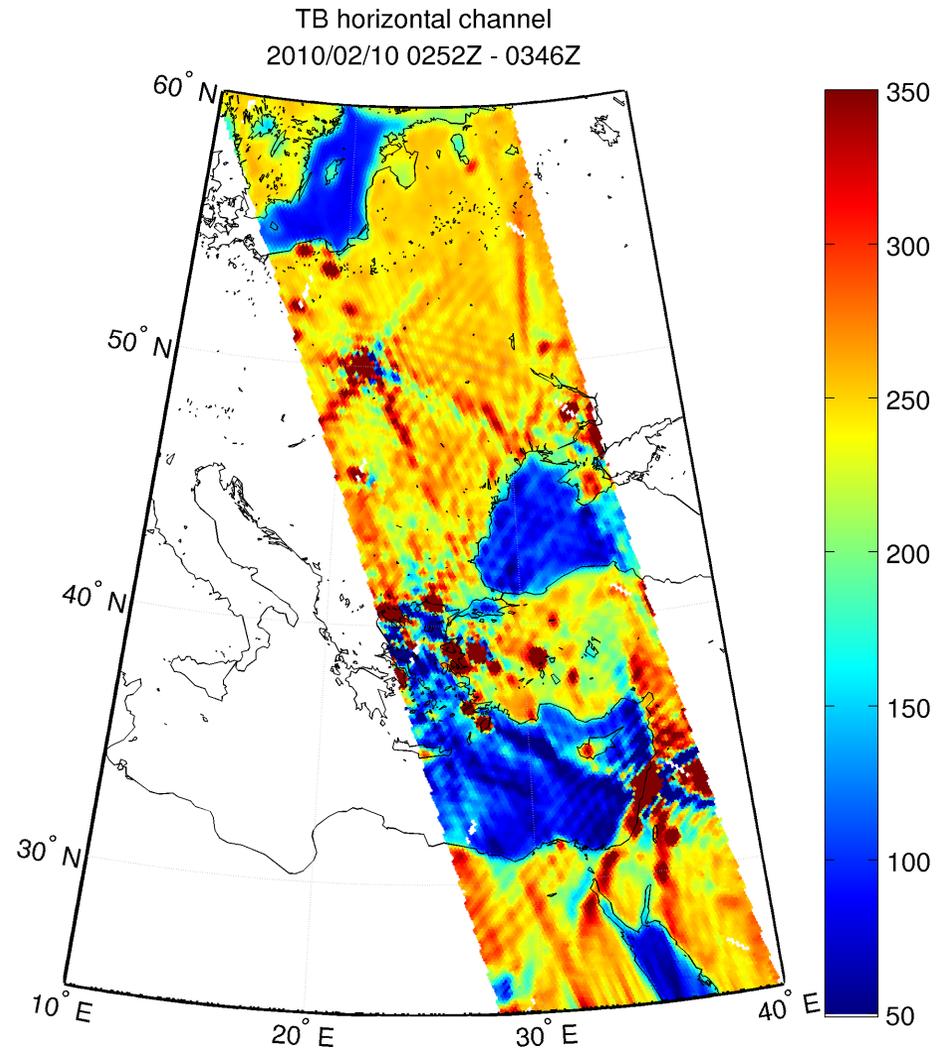


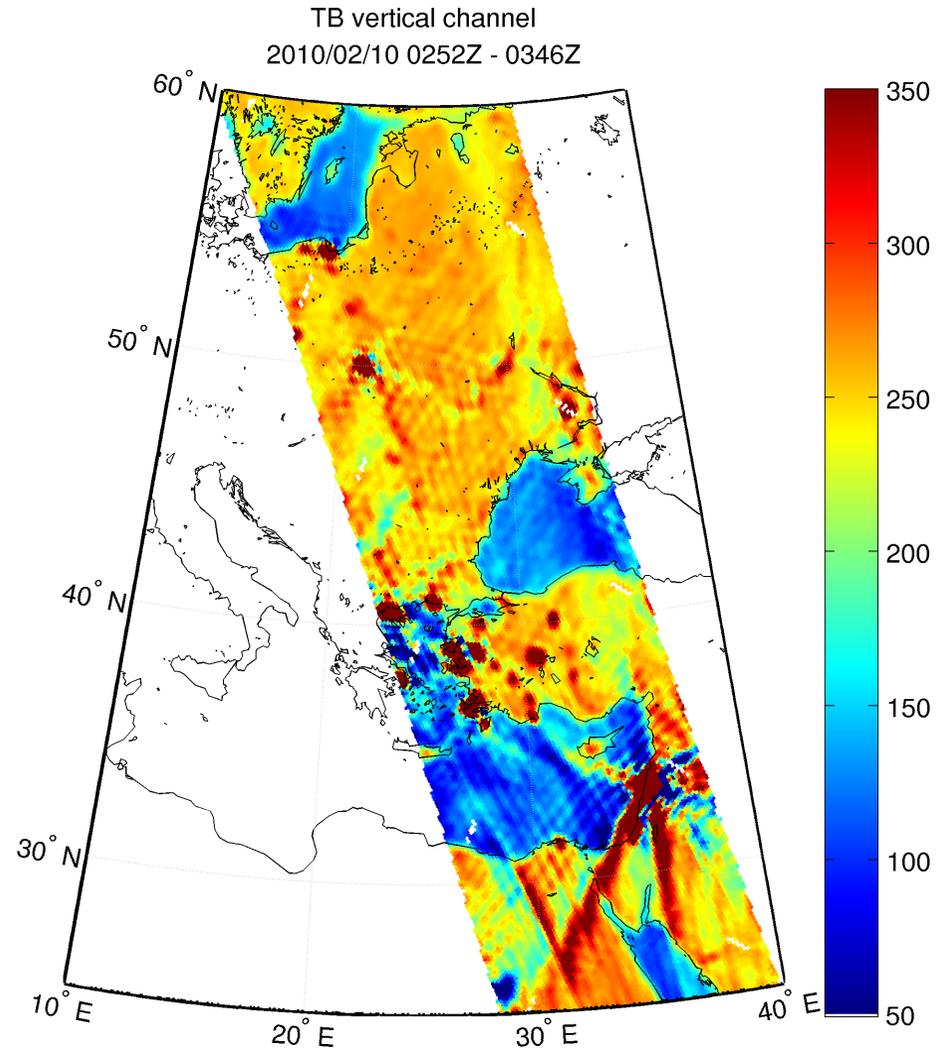
Conclusions From Airborne Measurements

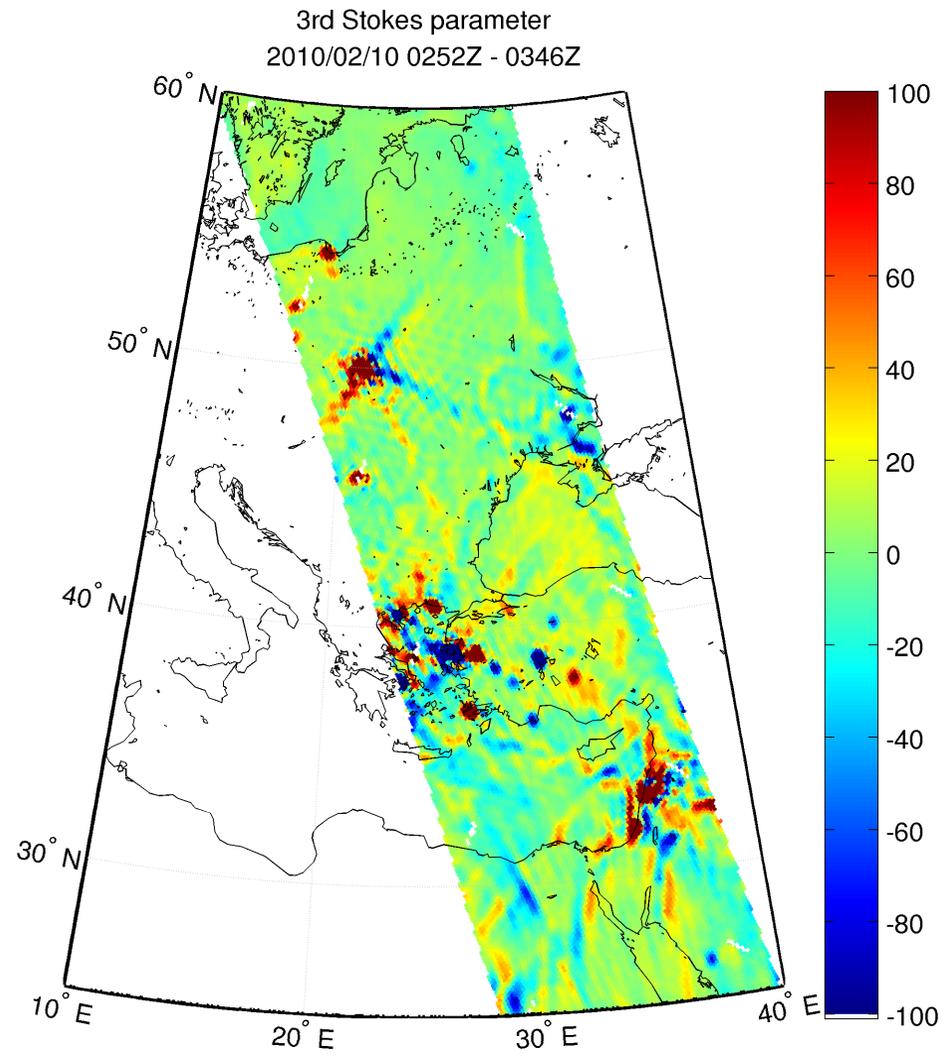
- **RFI is generally there**
- RFI is very variable in nature, time, and space
- DTU flights have generally experienced 1% RFI flagging over rural research areas - no problem for science data when you know
- **Kurtosis flagging often works quite well**
- Some additional cleaning is often needed

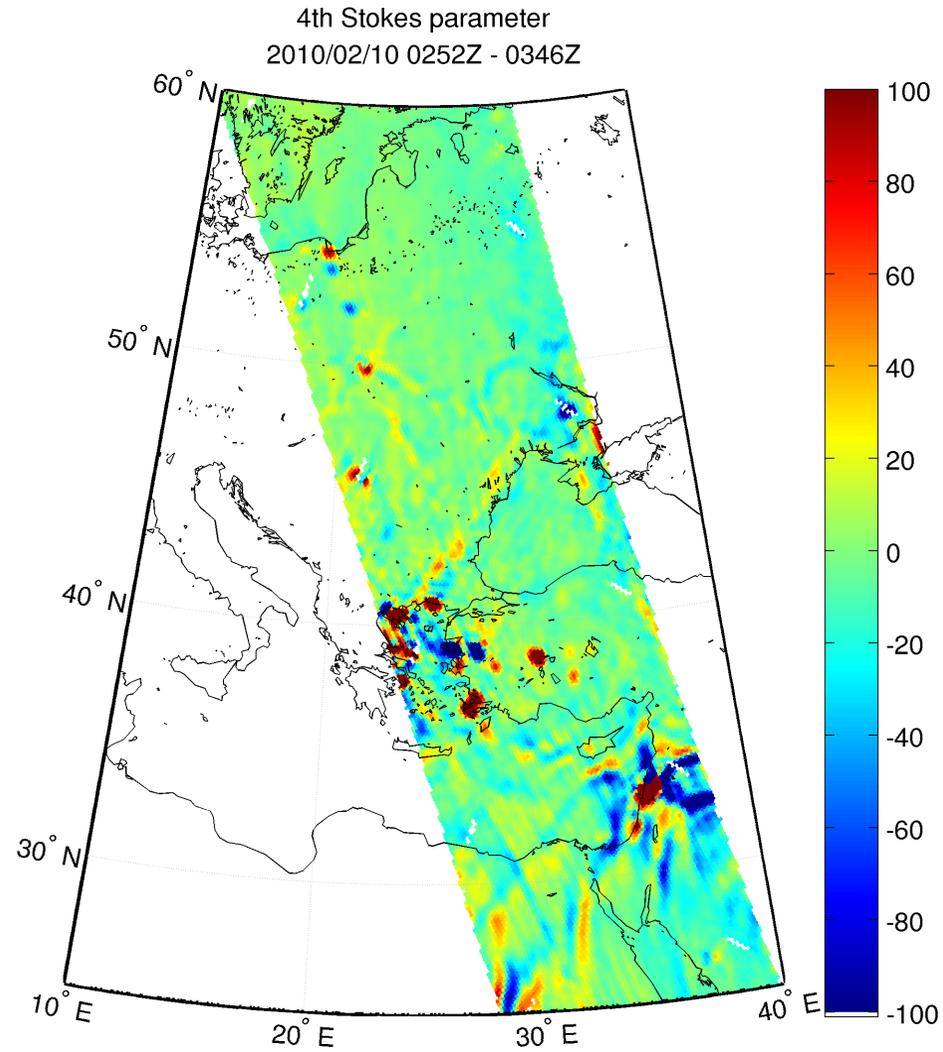
- Signatures in the 3rd and 4th Stokes channels indicate RFI
- A combination of kurtosis and the above is a powerful tool.
- But also polarimetry alone seems to work well
 - especially using the standard deviation of the 3rd and 4th Stokes parameters as indicators

- **Glitch detector cleans up very well after polarimetry has been used!**







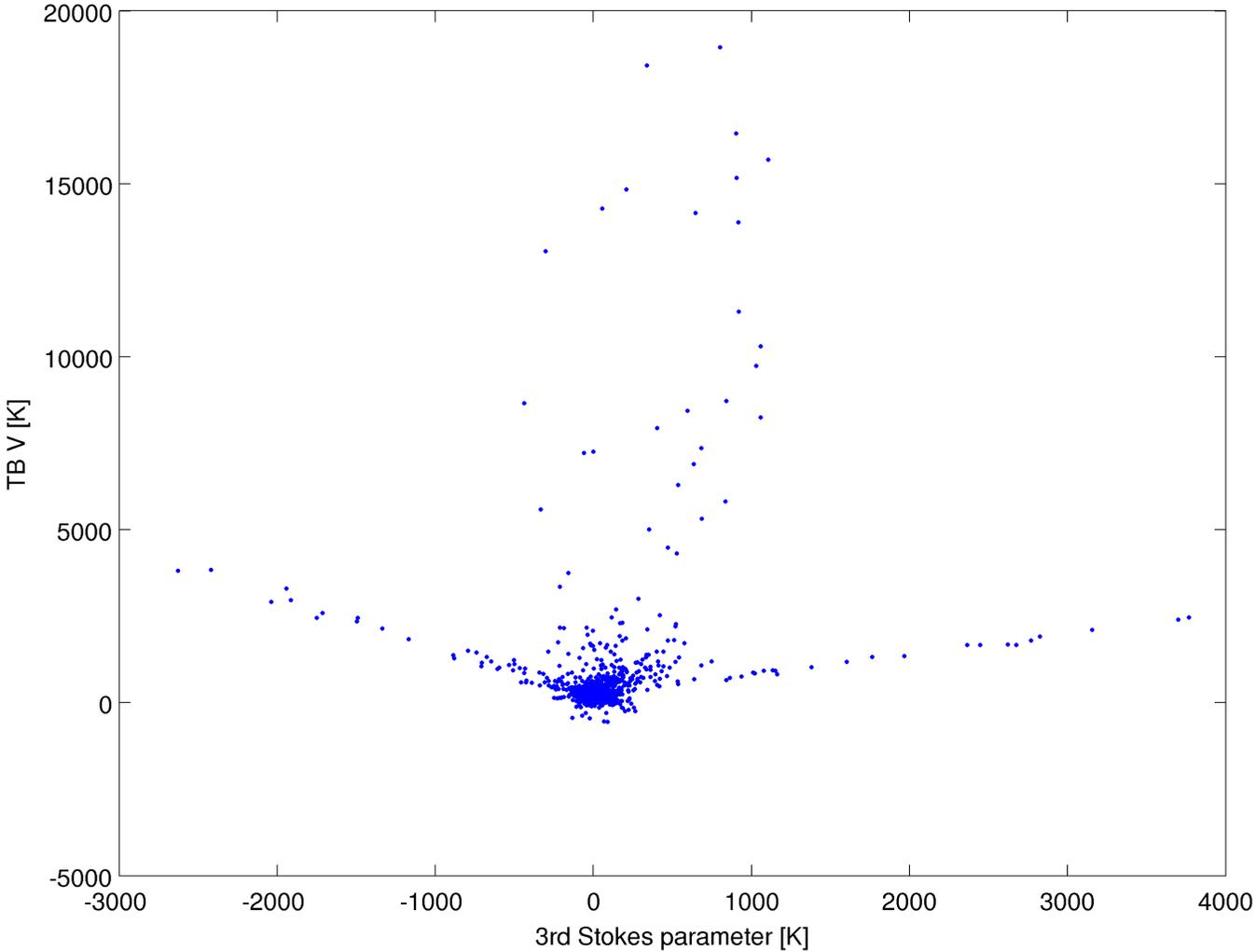


SMOS Polarimetric Mode

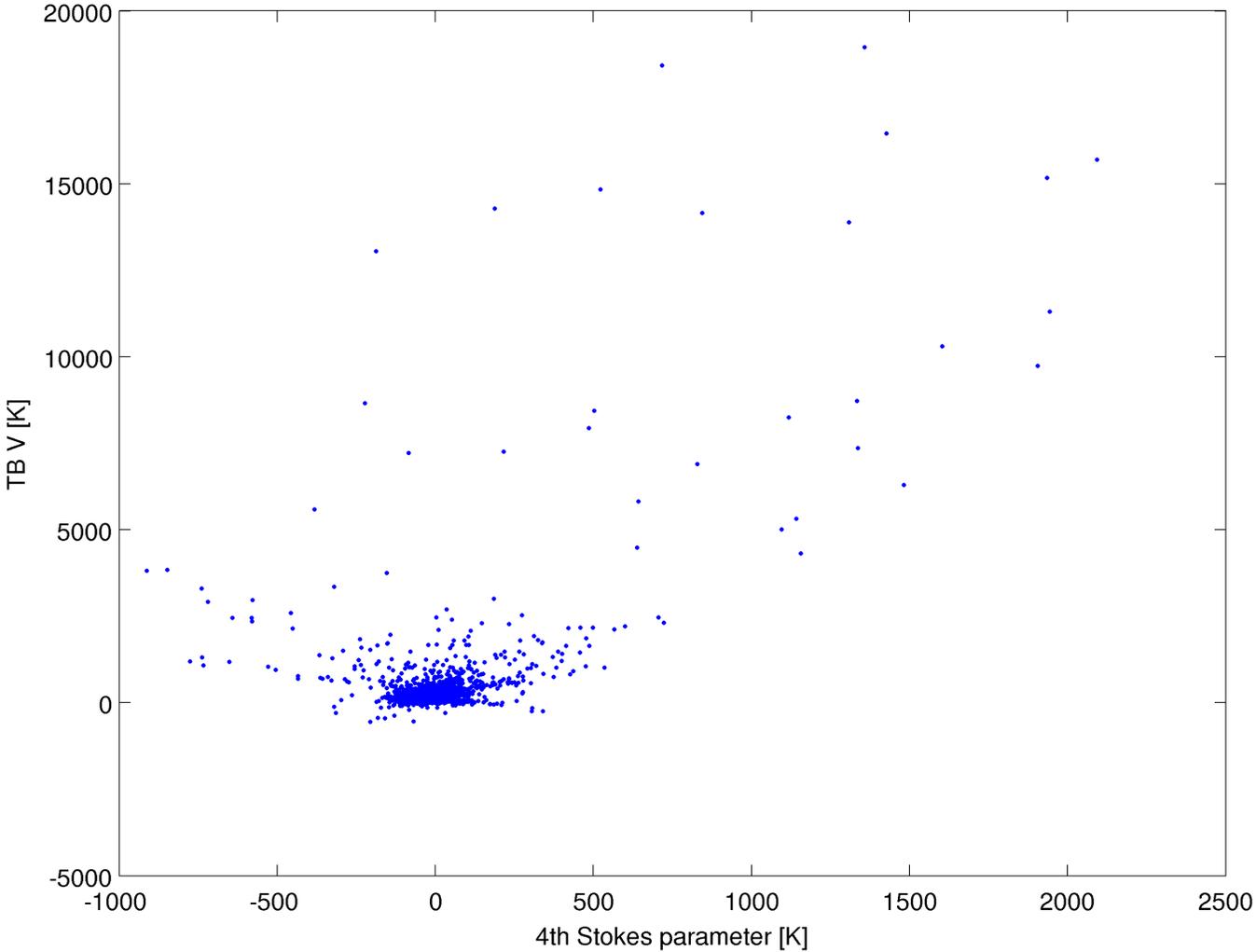
Mode:	1	2	3	4	1'	2'	3'	4'
Time:	1 τ		2 τ		3 τ		4 τ	
Result:	VV	HH/HV/VH		HH	VV/VH/HV			

SMOS MEASURES DIFFERENT POLARIZATIONS SEQUENTIALLY

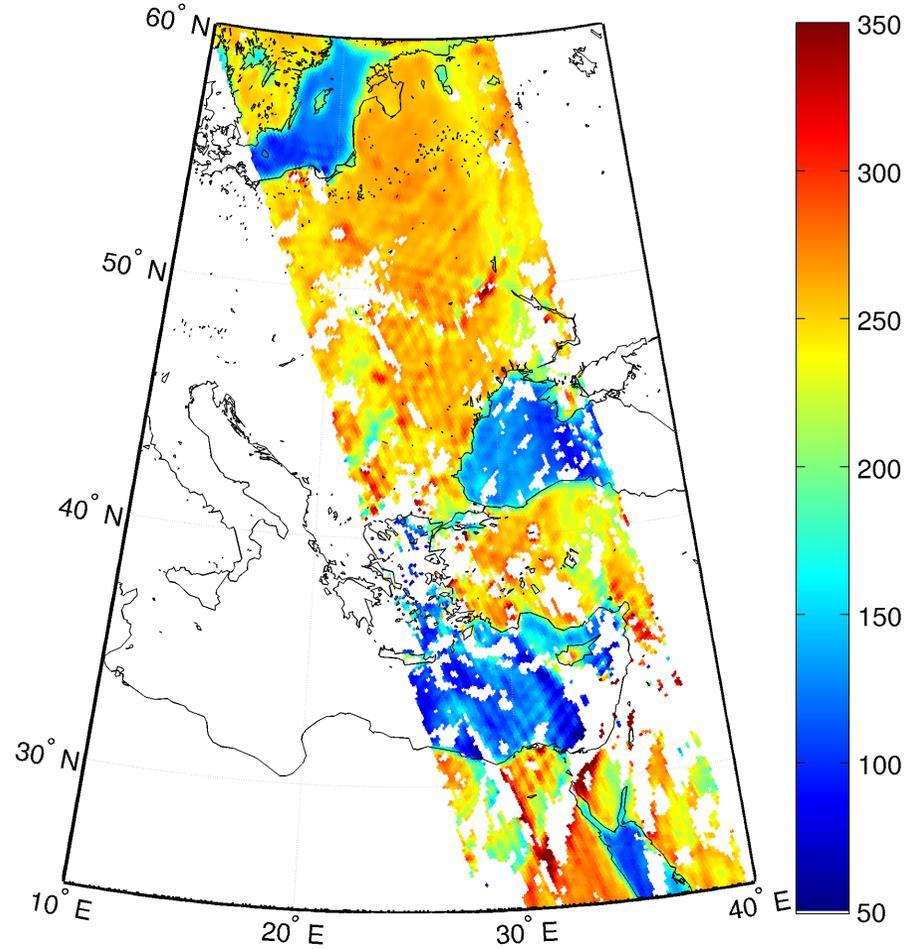
TB V vs. 3rd Stokes parameter
2010/02/10 0252Z - 0346Z



TB V vs. 4th Stokes parameter
2010/02/10 0252Z - 0346Z



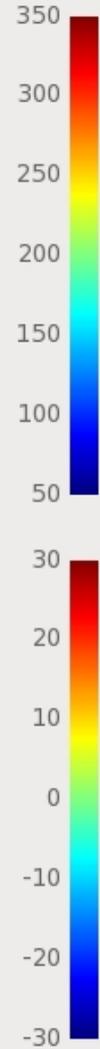
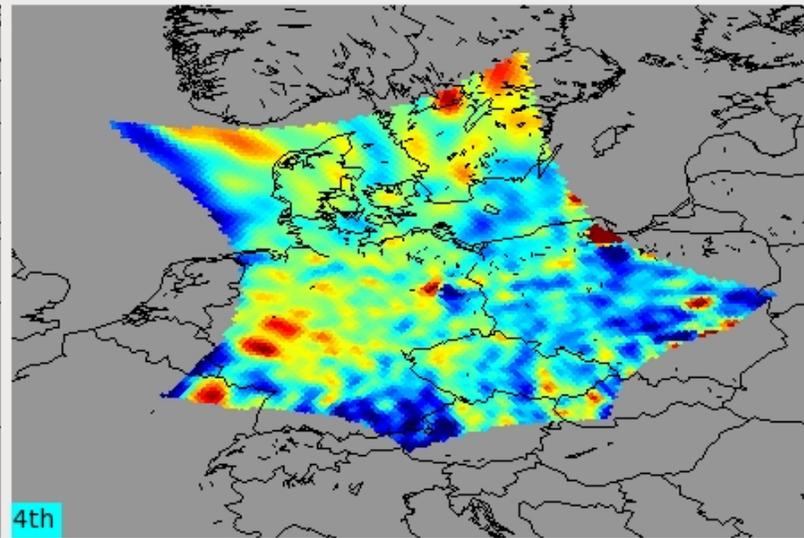
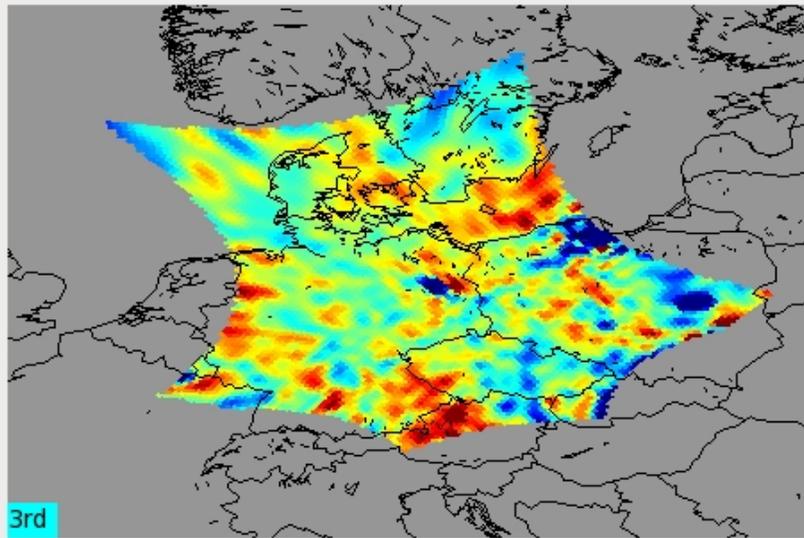
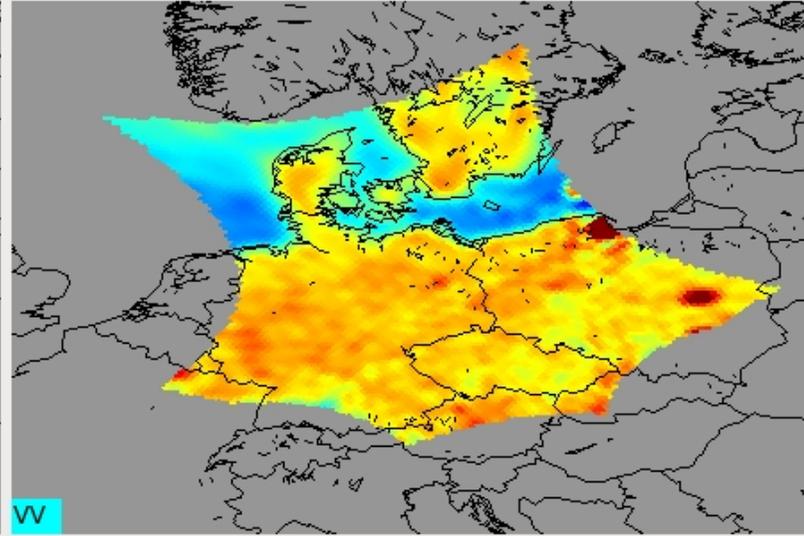
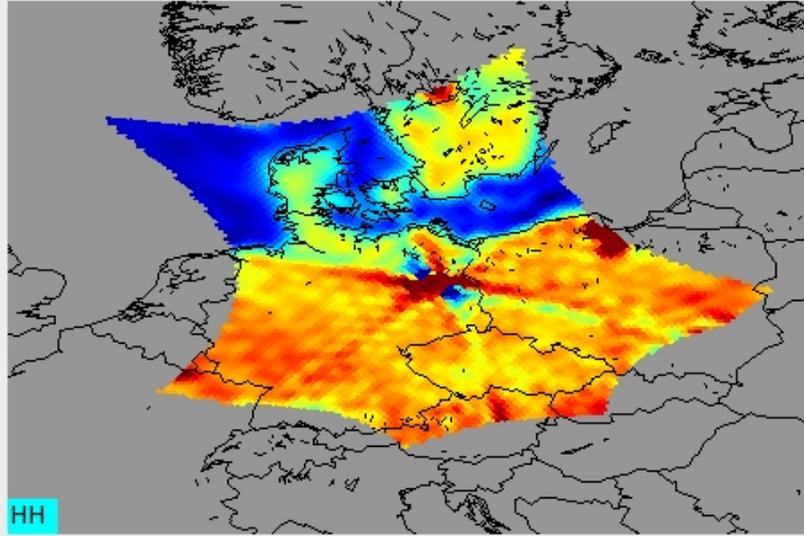
TB vertical channel
Cleaned by interting NaNs
2010/02/10 0252Z - 0346Z



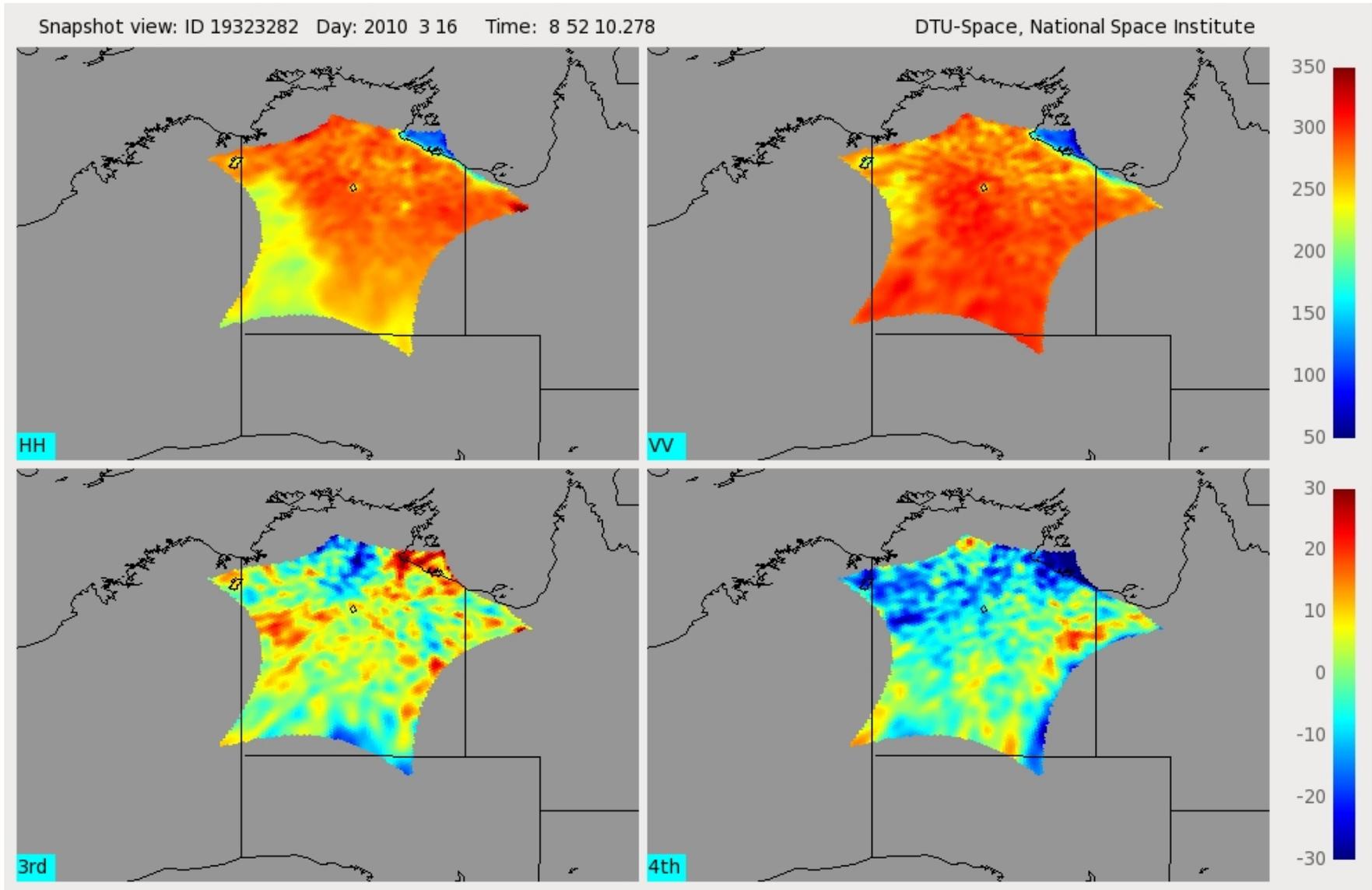
SMOS View of Europe

Snapshot view: ID 26060838 Day: 2010 5 2 Time: 4 21 44.546

DTU-Space, National Space Institute



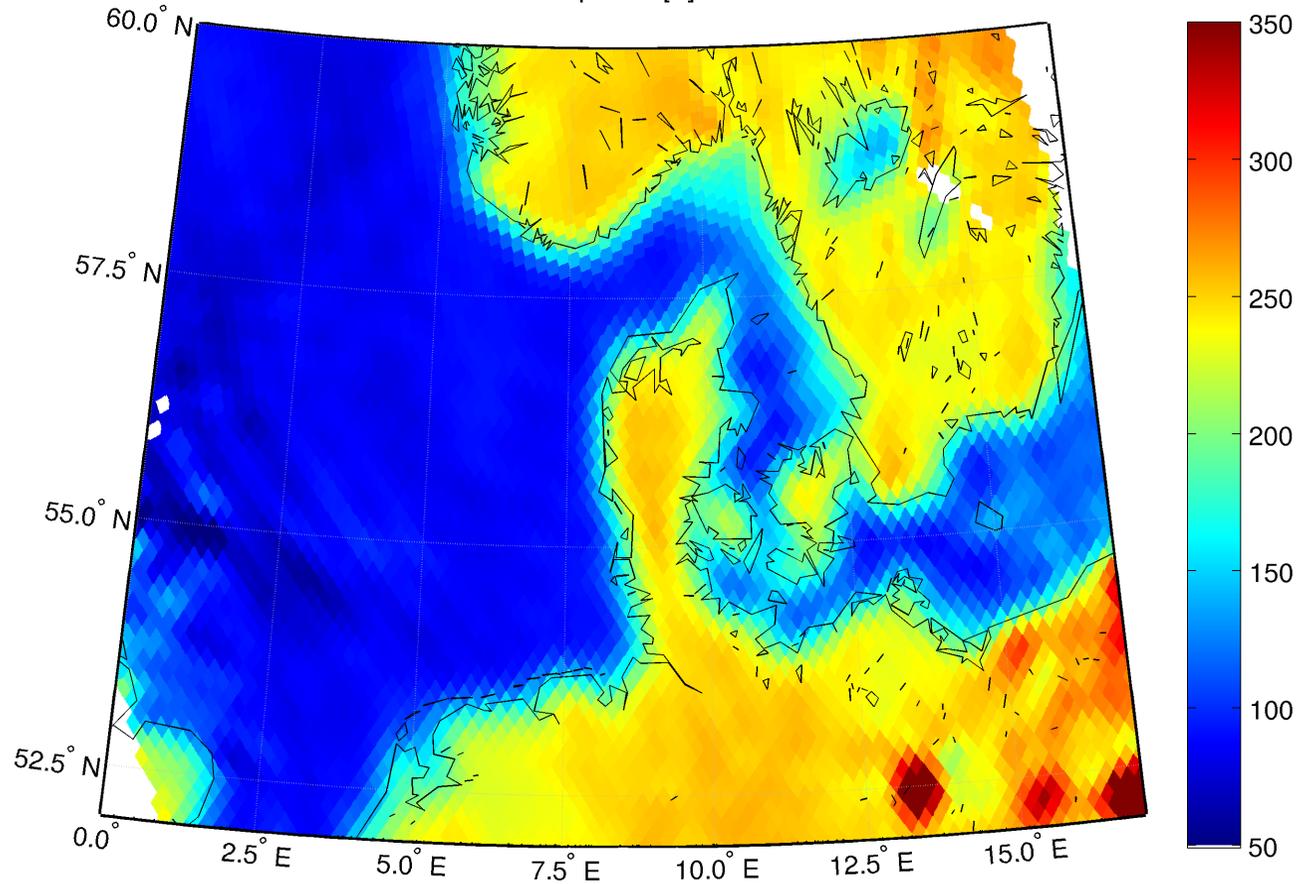
SMOS View of Australia



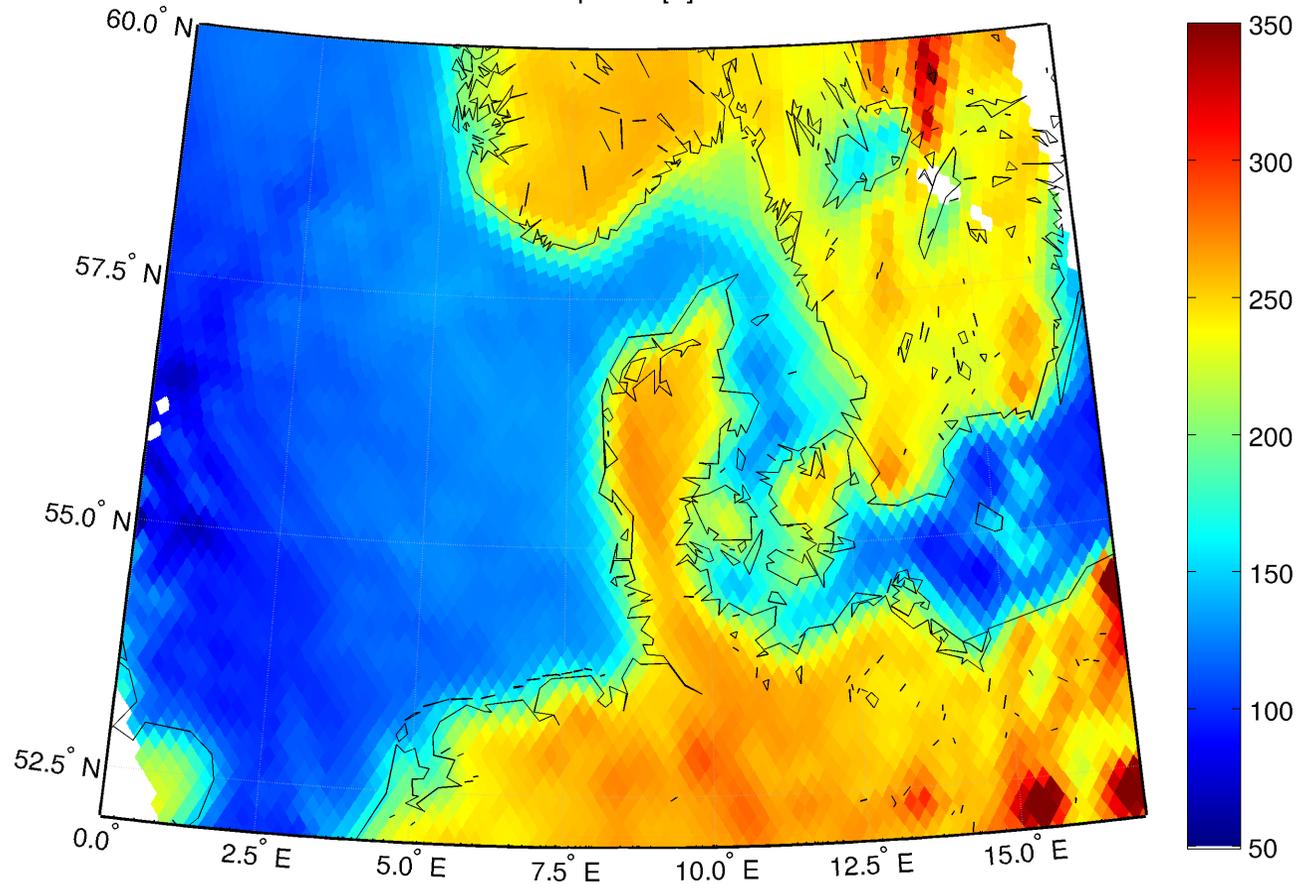
Conclusion

- **RFI is generally there**
- **RFI is very variable in time, and space**
- **Polarimetric signatures from SMOS looks like airborne ditto!**
- hence methods investigated using airborne data can be applied.
- **Polarimetry looks promising**
- **Glitch detector after polarimetry looks very promising**

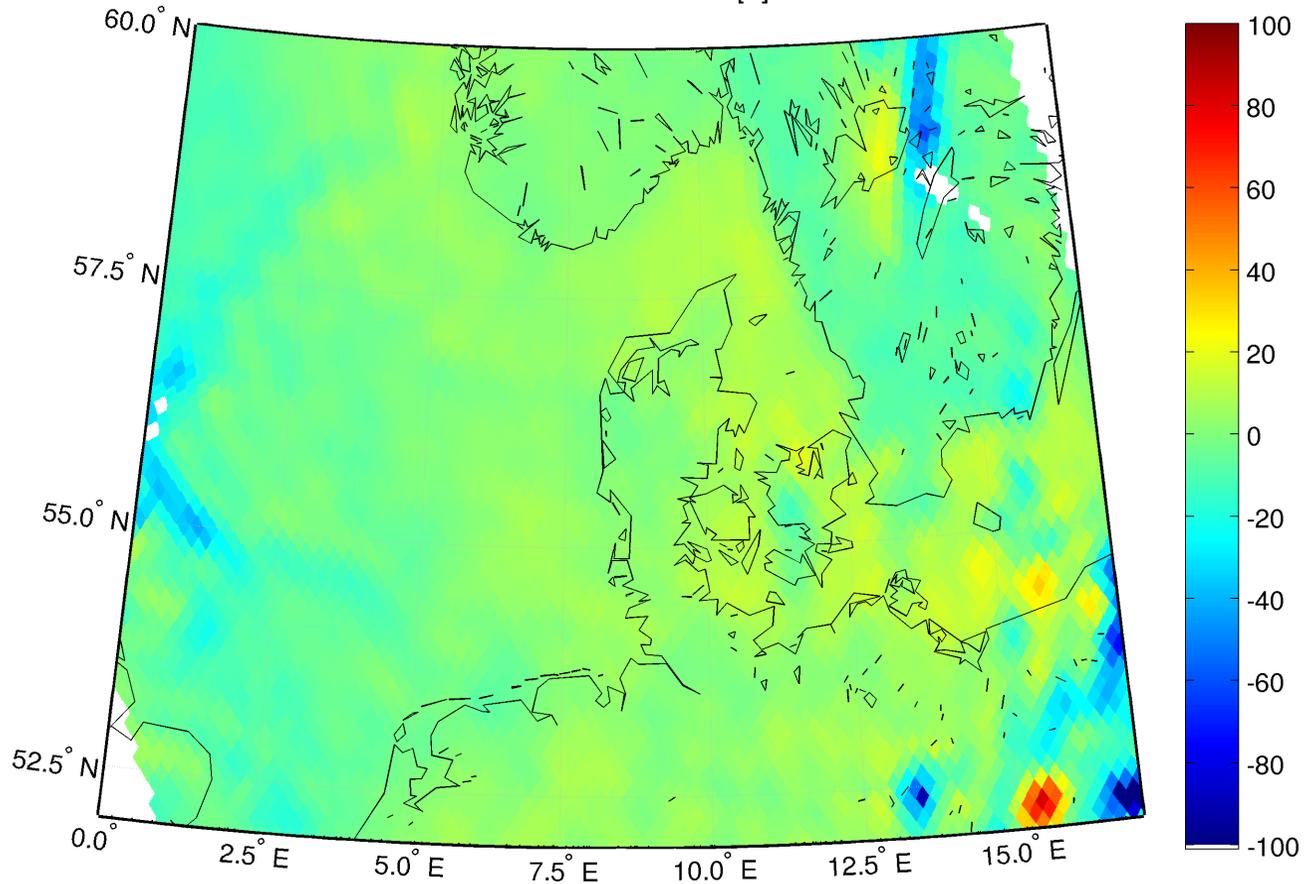
SMOS browse product
Swath started @ 20100211T035311, ended @ 20100211T044712
H-pol TB [K]



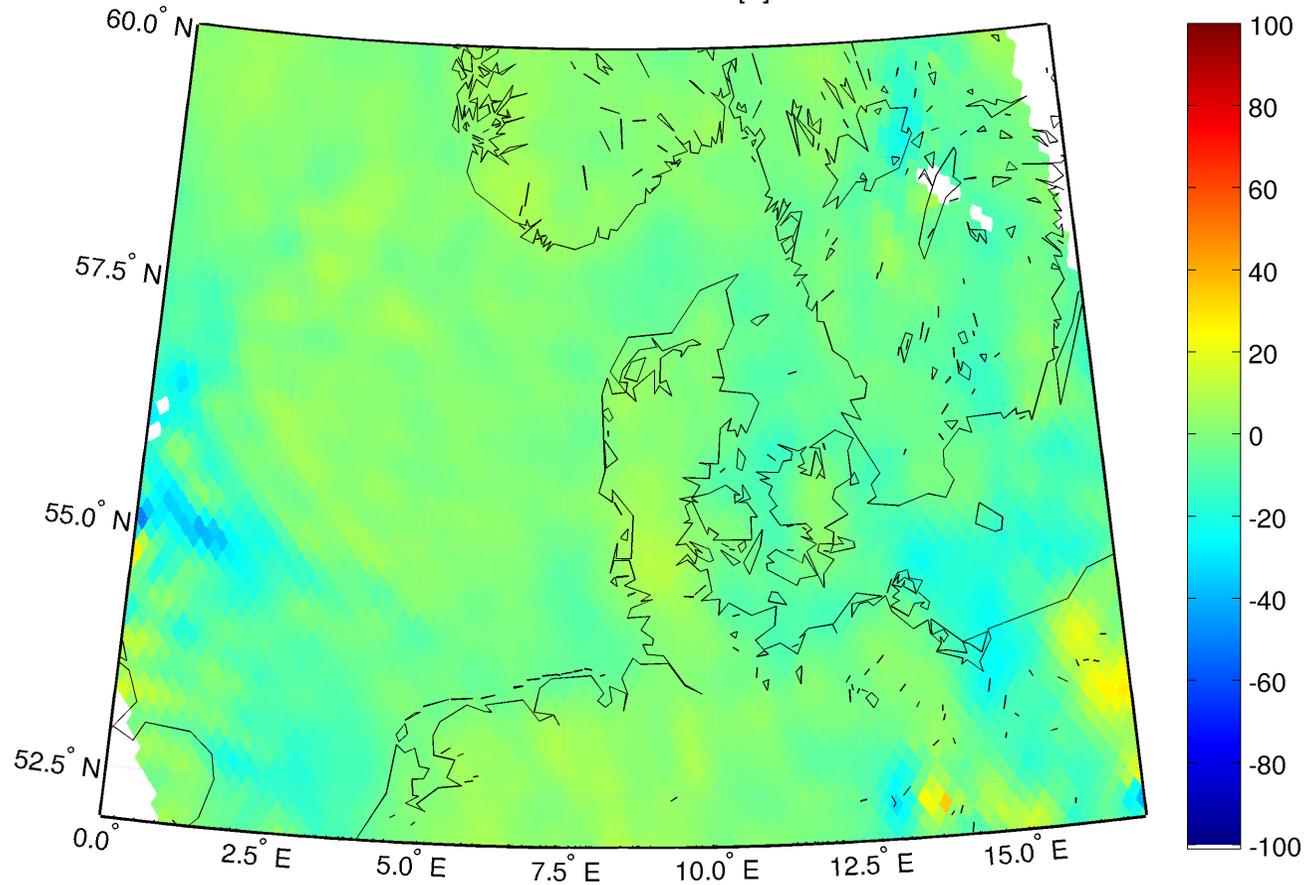
SMOS browse product
Swath started @ 20100211T035311, ended @ 20100211T044712
V-pol TB [K]



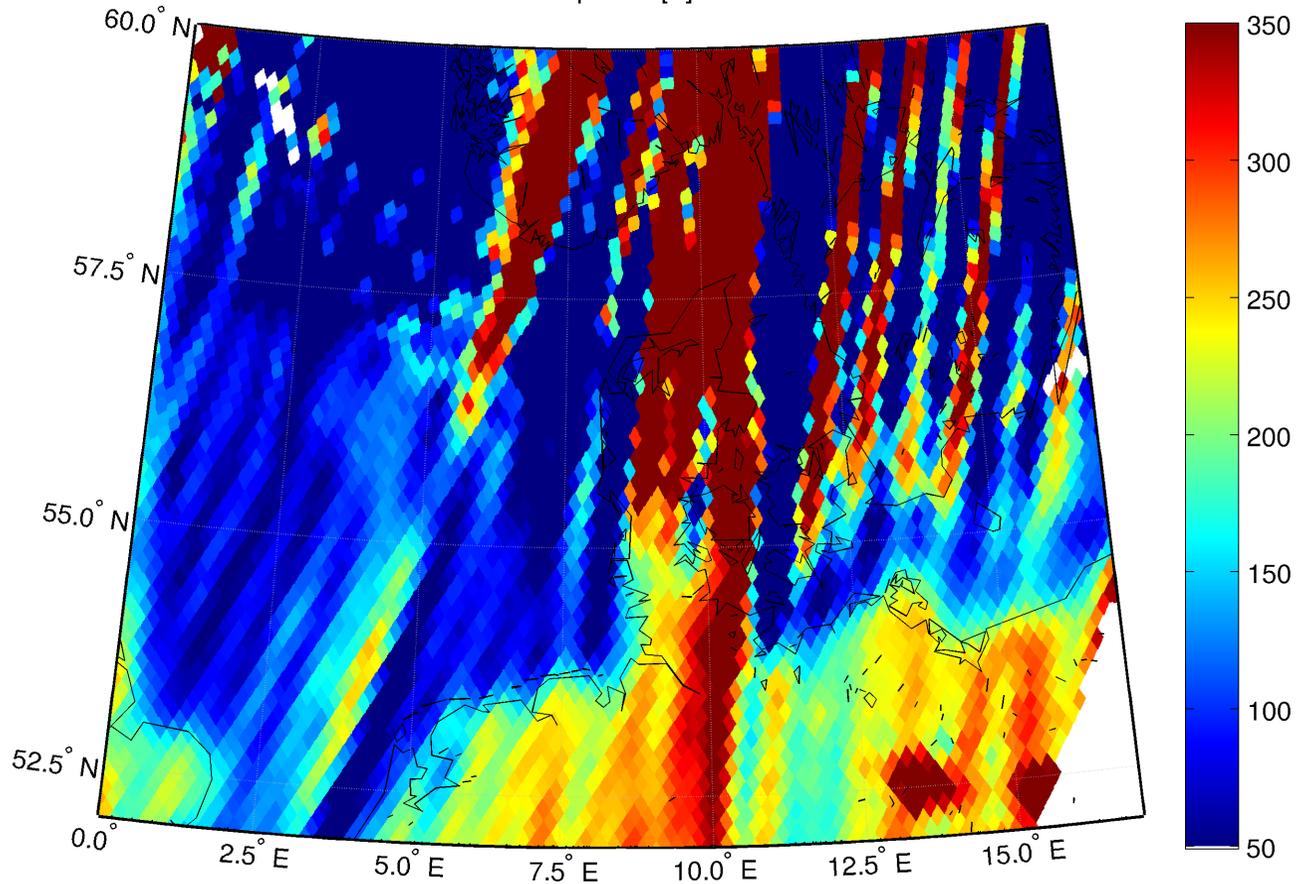
SMOS browse product
Swath started @ 20100211T035311, ended @ 20100211T044712
3rd Stokes Parameter [K]



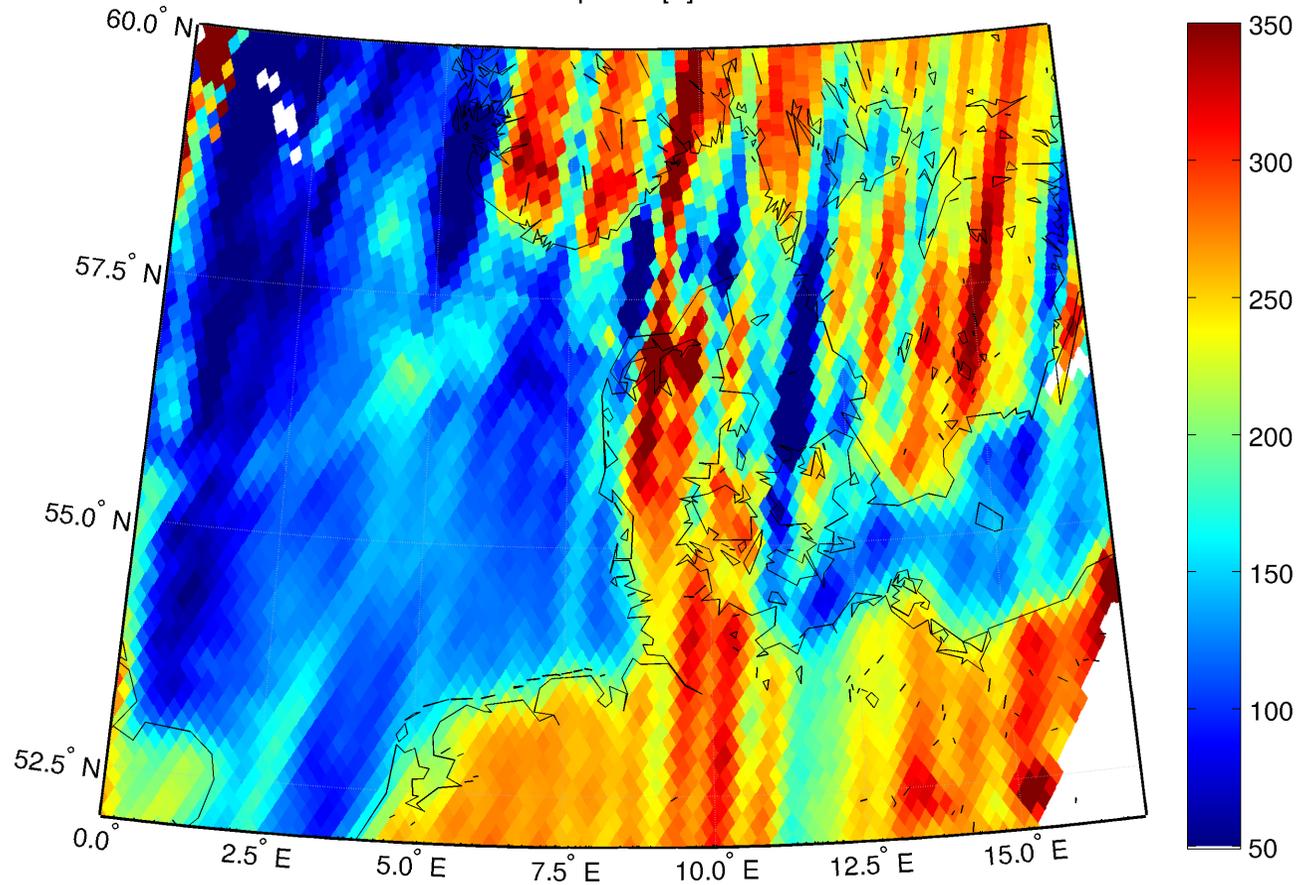
SMOS browse product
Swath started @ 20100211T035311, ended @ 20100211T044712
4th Stokes Parameter [K]



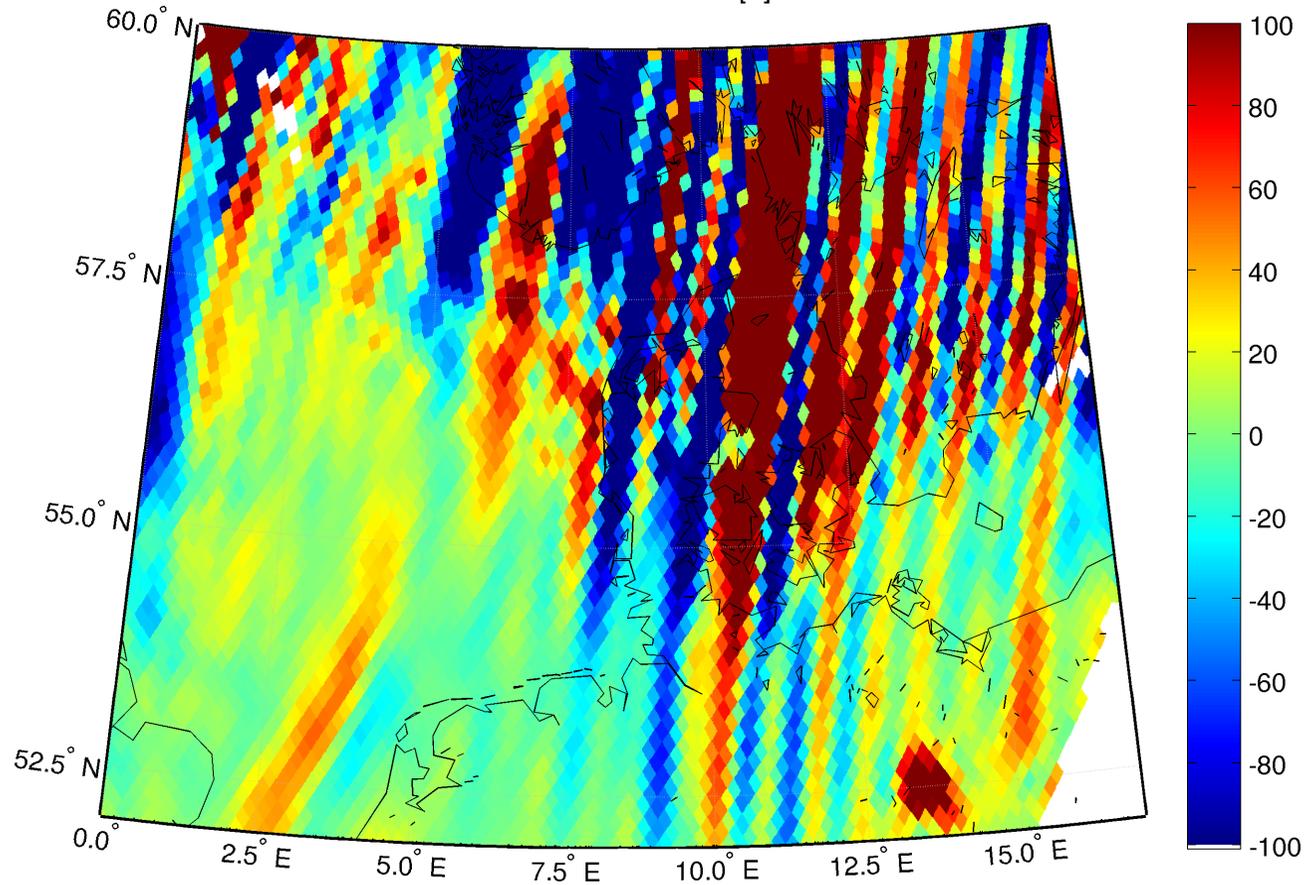
SMOS browse product
Swath started @ 20100211T180350, ended @ 20100211T185749
H-pol TB [K]



SMOS browse product
Swath started @ 20100211T180350, ended @ 20100211T185749
V-pol TB [K]



SMOS browse product
Swath started @ 20100211T180350, ended @ 20100211T185749
3rd Stokes Parameter [K]



SMOS browse product
Swath started @ 20100211T180350, ended @ 20100211T185749
4th Stokes Parameter [K]

